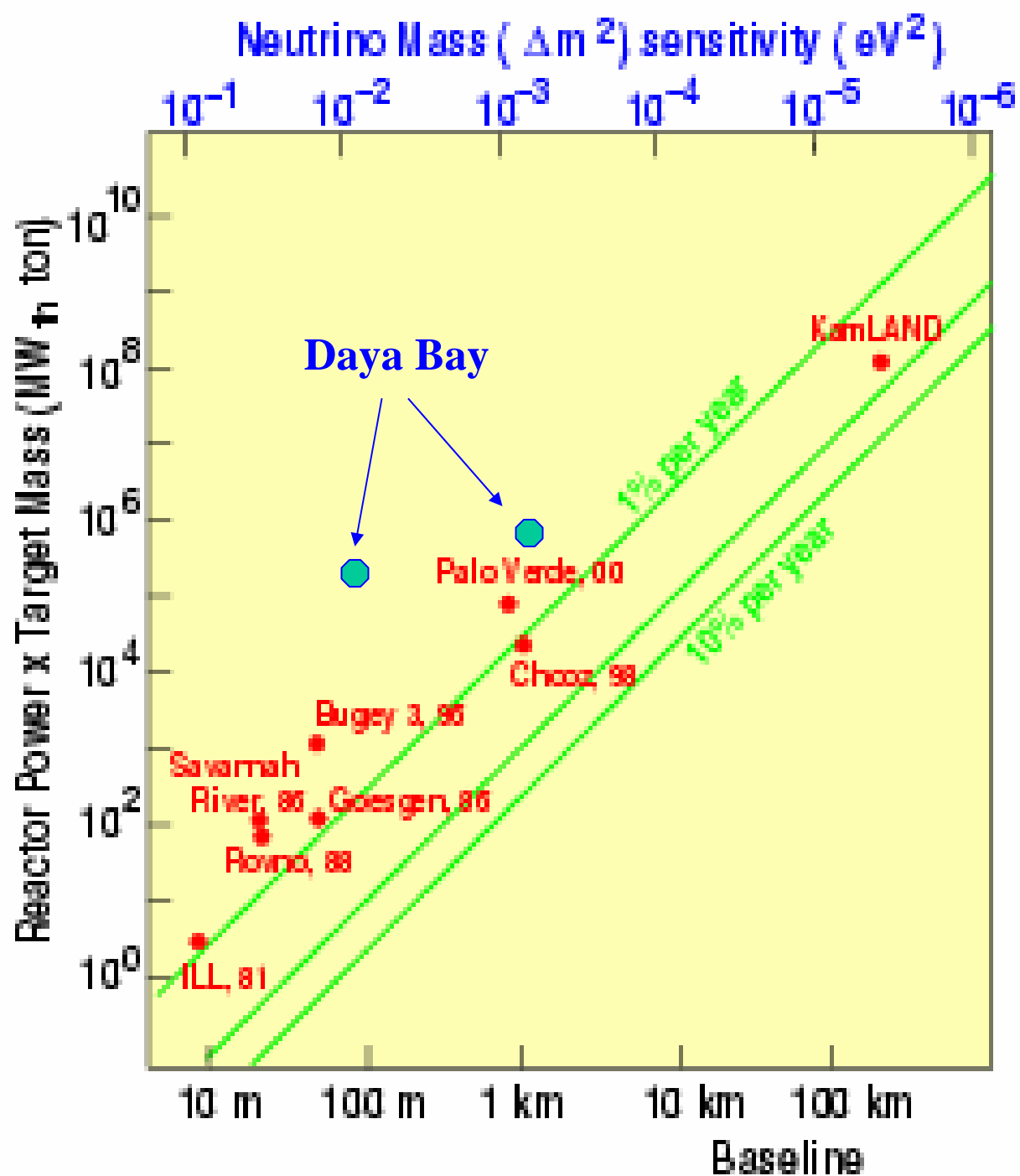


# **Systematic errors of reactor neutrino experiments**

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**Jan. 17, 2004**



# Systematic errors

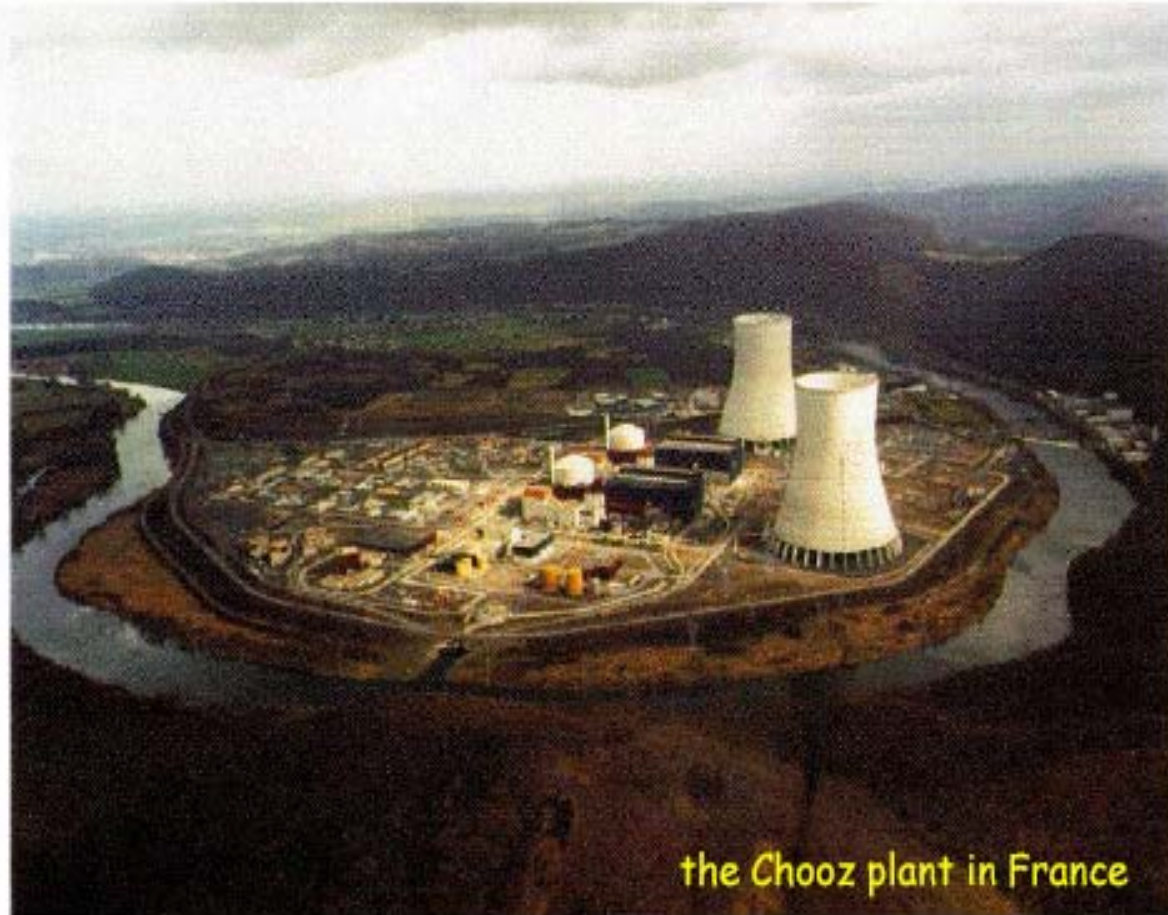
- Reactor related error
- Detector related error
- Background related error

# Reactor: Source of neutrinos

$200\text{MeV} / \text{fission}$

$6\bar{\nu}_e / \text{fission}$

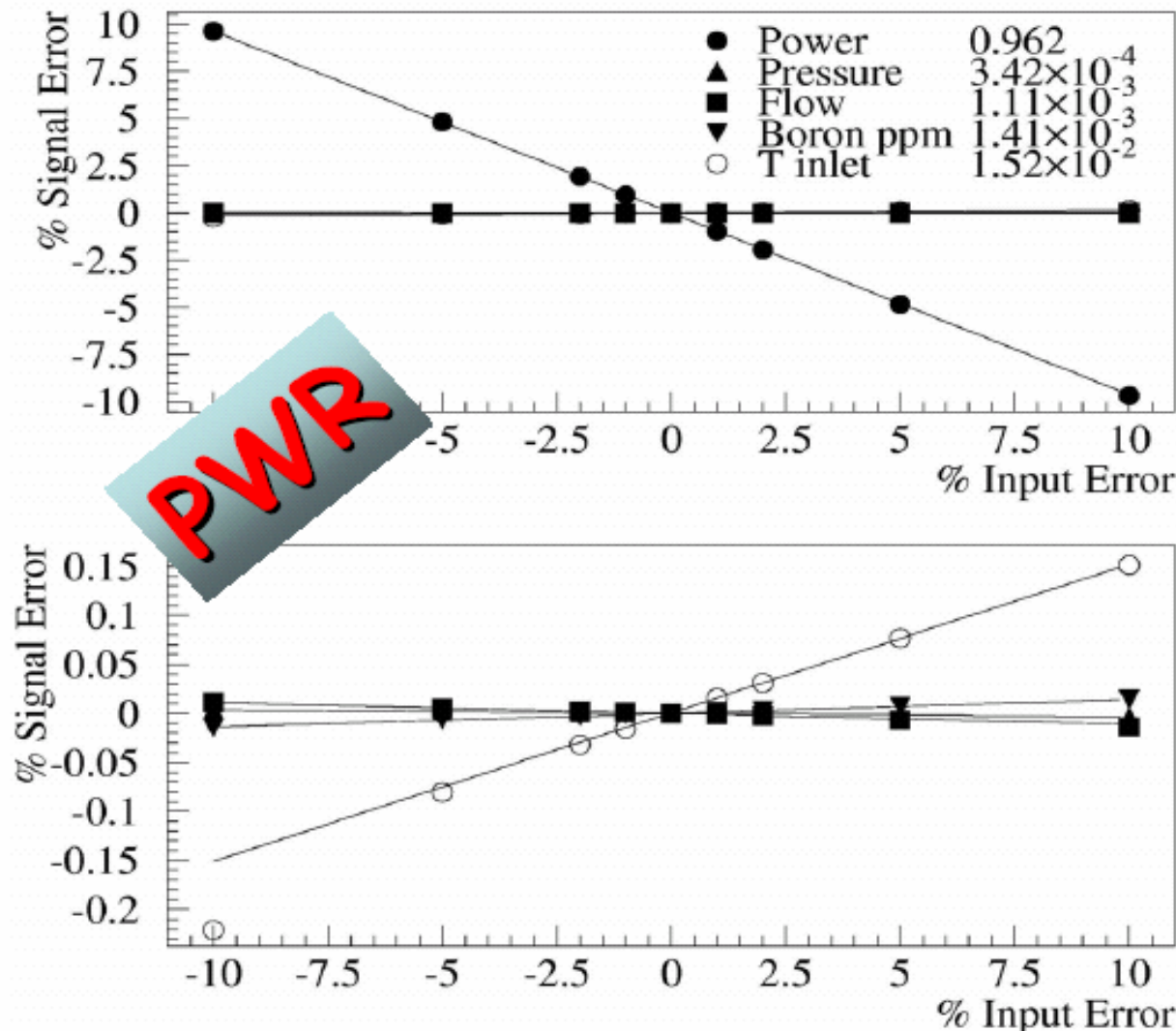
A typical large power  
reactor produces  
 $3\text{ GW}_{\text{thermal}}$  and  
 $6 \cdot 10^{20}$  antineutrinos/s



the Chooz plant in France

# Systematic Error: Power

*The 200 MeV/fission part:*



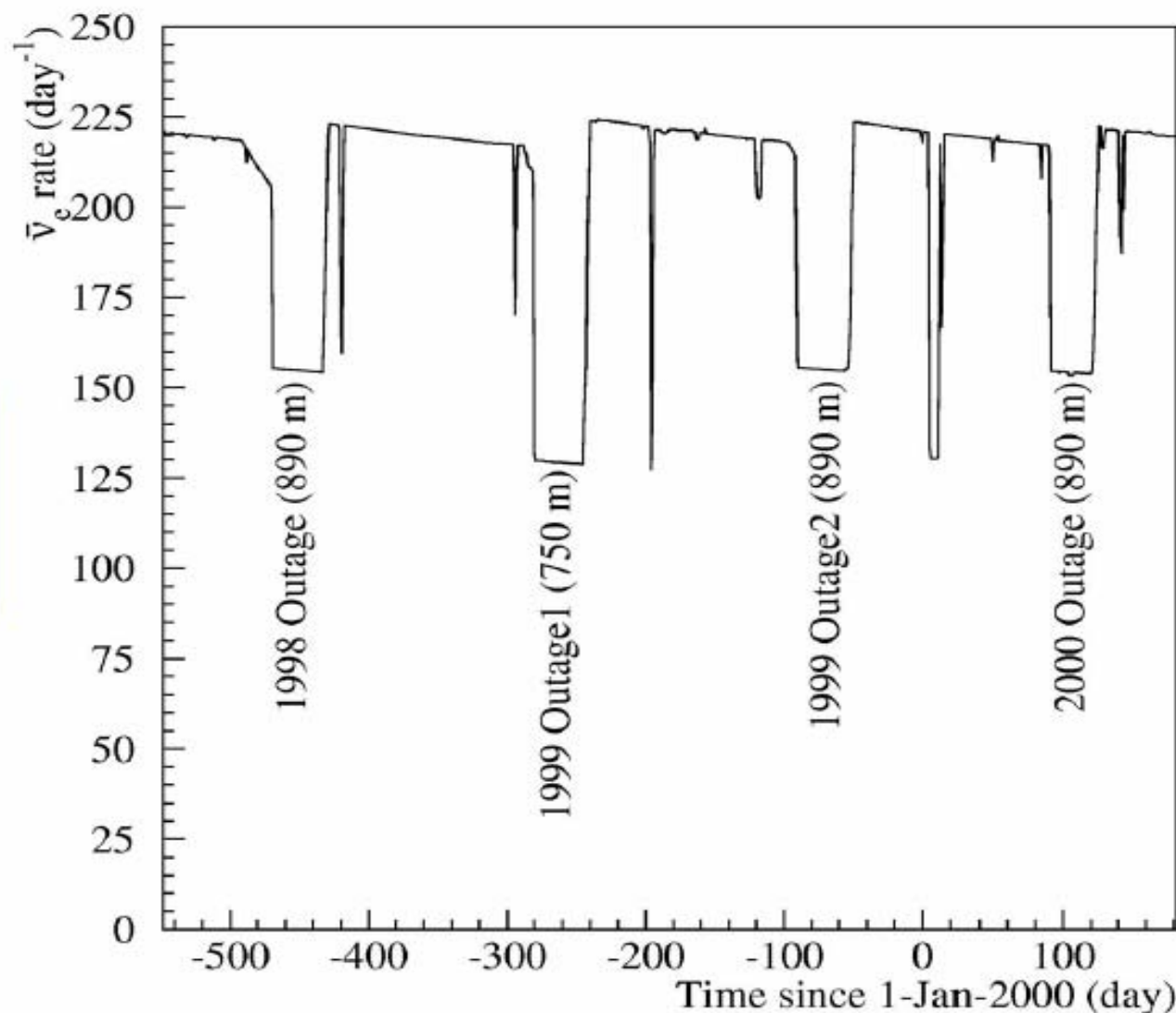
Thermal power is routinely measured by the reactor operator in order to adjust the reactor to the highest licensed power

*Economics push the error on this to 0.6-0.7%*

# Reactor thermal power

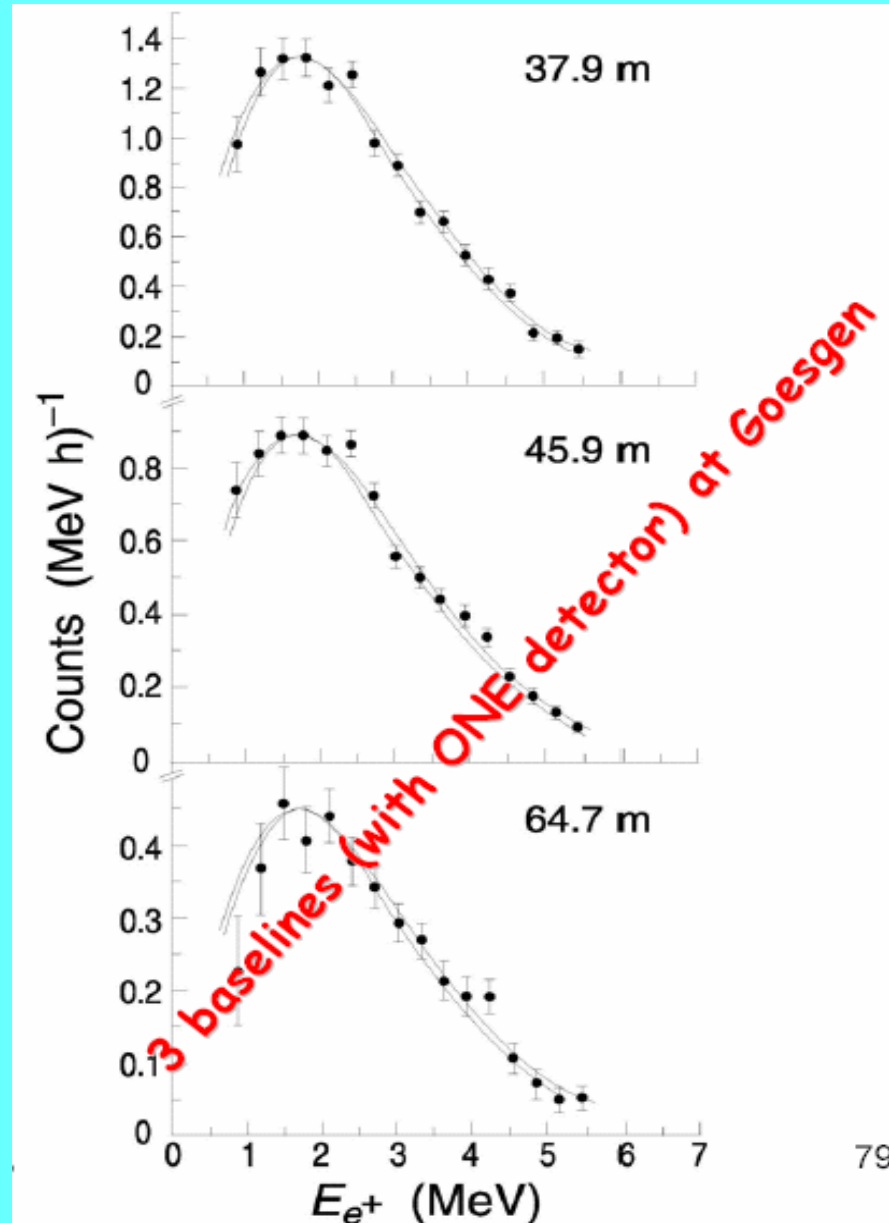
Known to <1%

Power history  
for the 3-unit  
Palo Verde plant



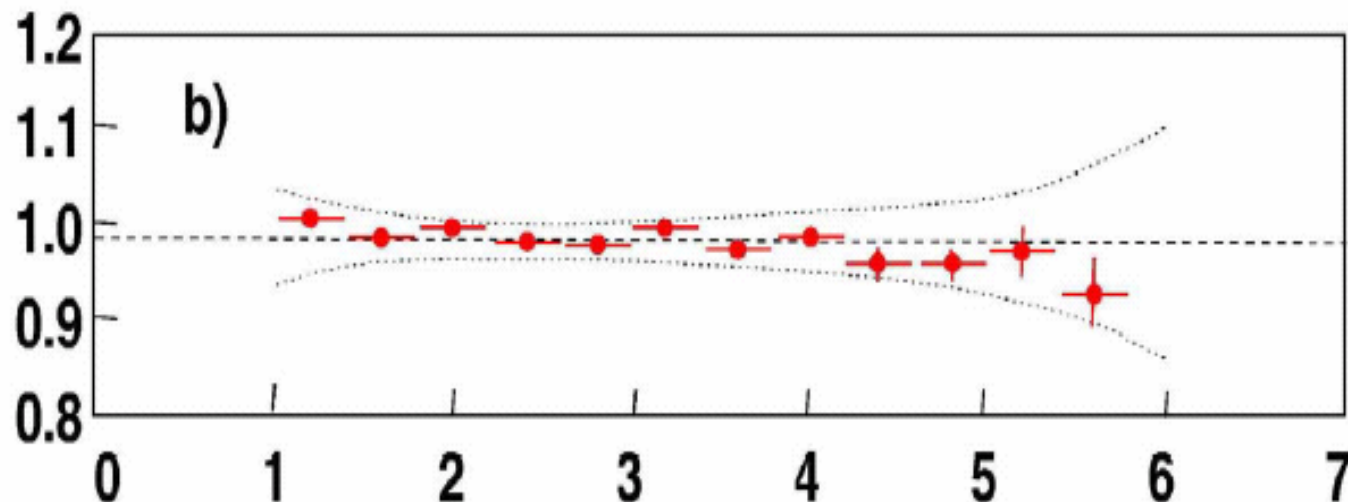
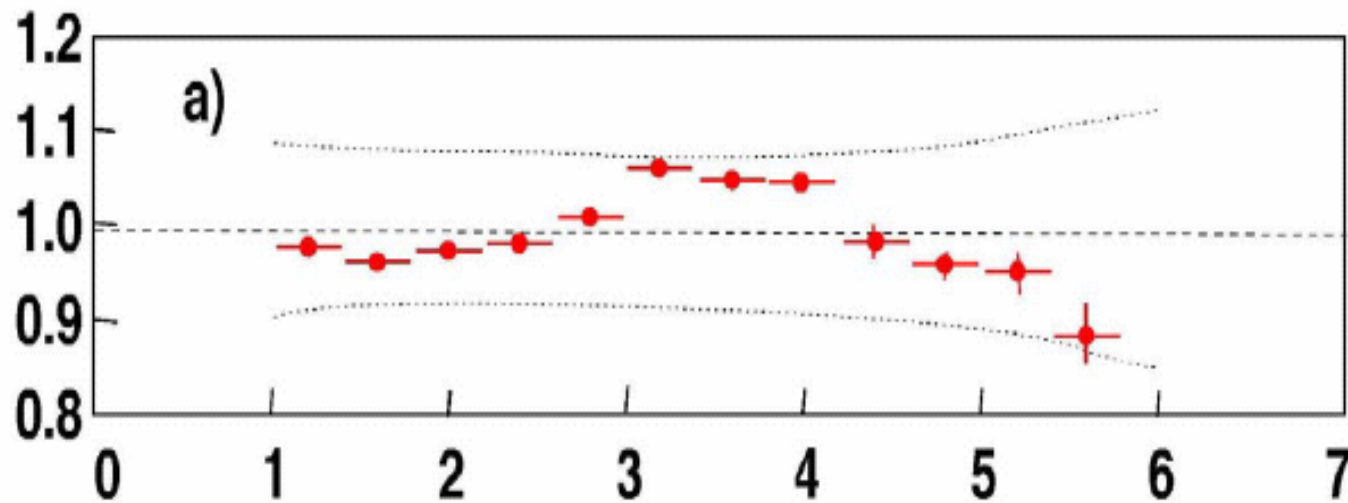
# Prediction of reactor neutrino spectrum

- Three ways to obtain reactor neutrino spectrum:
  - Direct measurement
  - First principle calculation
  - Sum up neutrino spectra from  $^{235}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$  and  $^{238}\text{U}$   
 $^{235}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$  from their measured  $\beta$  spectra  
 $^{238}\text{U}$ (10%) from calculation (10%)
- They all agree well within 3%





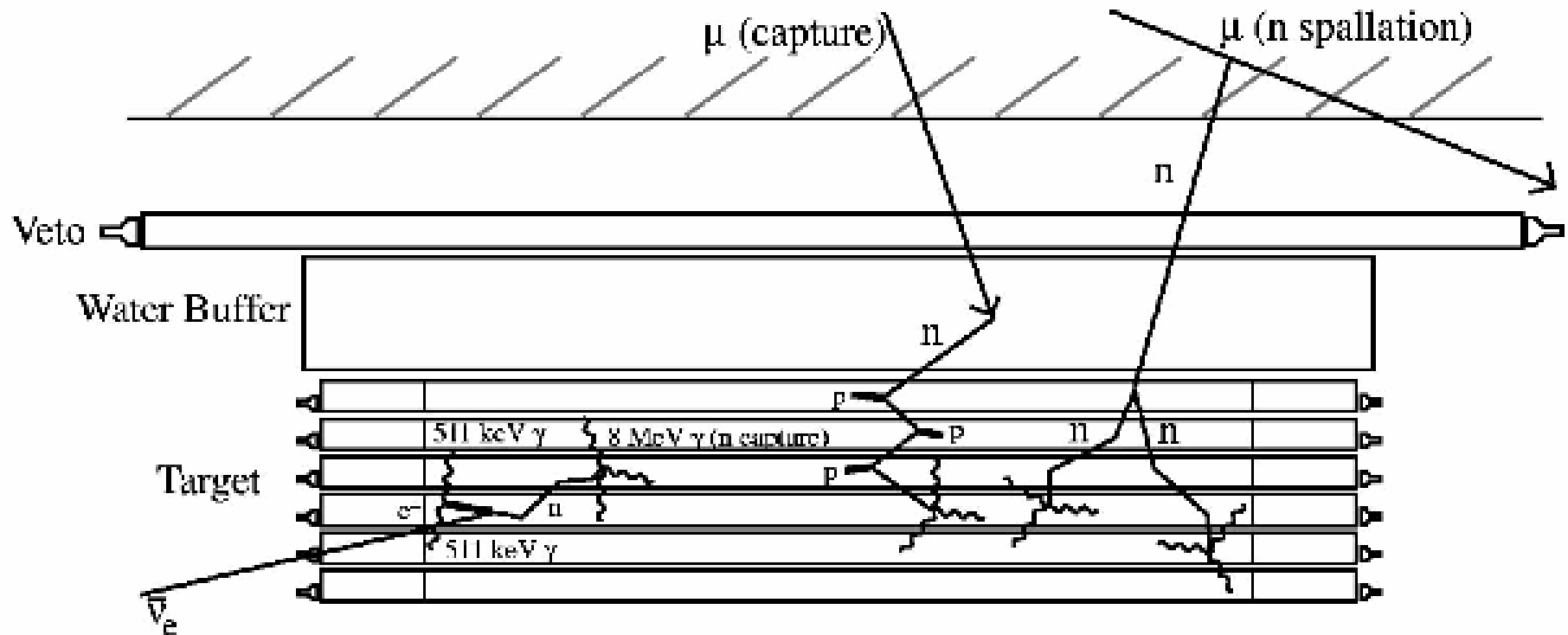
# Total error on neutrino spectrum



Positron energy (MeV)



# Background - Correlated



**Background - Uncorrelated: environmental radioactivity**

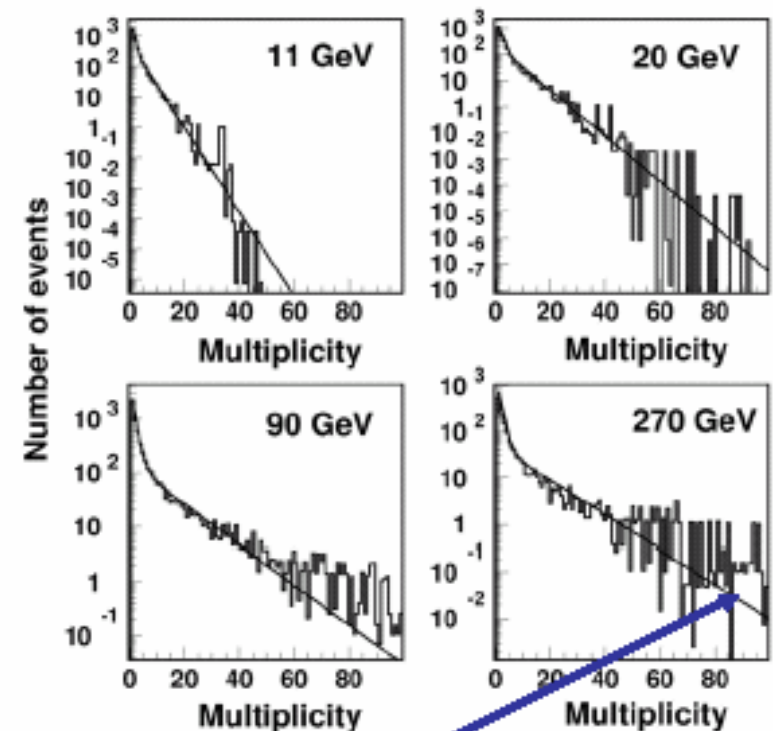
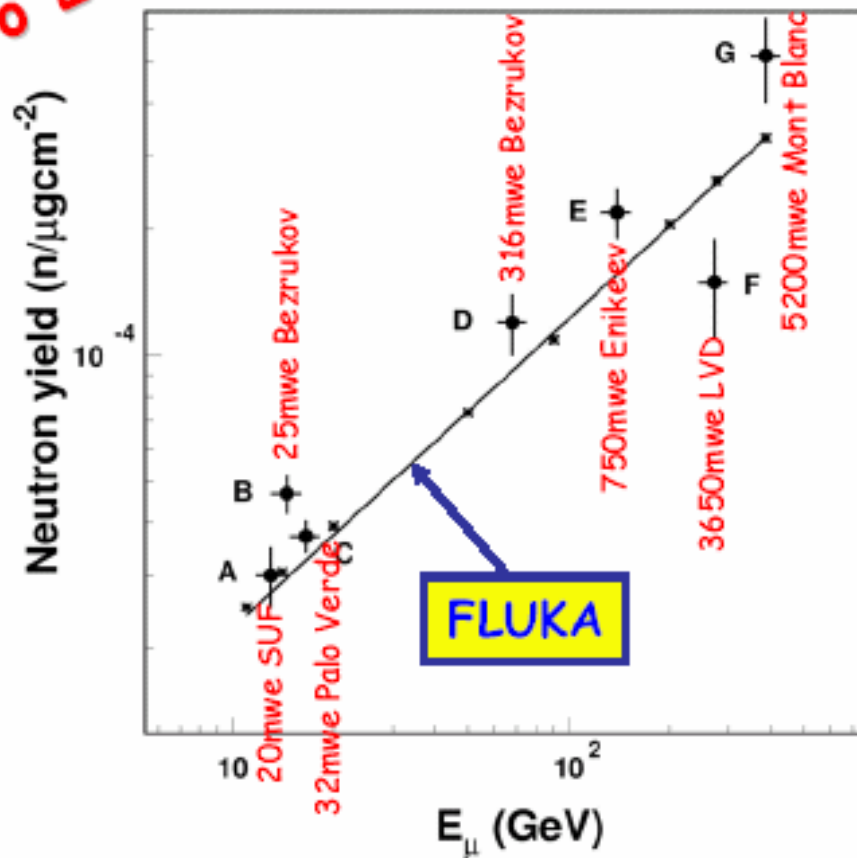
n-production from  $\mu$  quite well studied at Palo Verde, high energy spectrum and vetoed events can be used to predict background in the interesting region

Measurement of bkgnd to 20% appears possible

- Y-F. Wang, L. Miller, *GG; Phys. Rev. D* 62, 013012 (2000)
- F. Boehm et al., *Phys. Rev. D* 62, 092005 (2000)

FLUKA is quite accurate for most variables

- Y-F. Wang et al., *Phys. Rev. D* 64, 013012 (2001)



Parametrization based on FLUKA

Table 6

Muon-induced background rates in BOREXINO, calculated for different energy regions that are relevant for solar neutrino physics<sup>a</sup>

Isotopes	Muon-induced background rates in BOREXINO given in counts/(day $\times$ 100 tons) for the different energy regions			
	Full energy range	250 < $E$ < 800 keV <sup>7</sup> Be- $\nu$ region	0.8 < $E$ < 1.4 MeV pep- $\nu$ region	2.8 < $E$ < 5.5 MeV <sup>8</sup> B- $\nu$ region
<sup>11</sup> C	$14.55 \pm 1.49$	0	$7.36 \pm 0.75$	0
<sup>7</sup> Be	$0.34 \pm 0.04$	$0.34 \pm 0.04$	0	0
<sup>11</sup> Be	<0.034	$<4.3 \times 10^{-4}$	$<1.0 \times 10^{-4}$	<0.01
<sup>10</sup> C	$1.95 \pm 0.21$	0	0	$0.56 \pm 0.06$
<sup>8</sup> Li	$0.070 \pm 0.017$	$(2.5 \pm 0.6) \times 10^{-4}$	$(8.0 \pm 2.0) \times 10^{-4}$	$0.020 \pm 0.005$
<sup>6</sup> He	$0.26 \pm 0.03$	$0.040 \pm 0.004$	$0.07 \pm 0.01$	$0.011 \pm 0.001$
<sup>8</sup> B	$0.11 \pm 0.02$	0	$(3.3 \pm 0.6) \times 10^{-5}$	$0.020 \pm 0.004$
<sup>9</sup> C	$0.077 \pm 0.025$	0	0	$0.016 \pm 0.005$
<sup>9</sup> Li + <sup>8</sup> He	$0.034 \pm 0.007$	$<6.8 \times 10^{-4}$	$<1.0 \times 10^{-3}$	<0.014

<sup>a</sup> The rates are given in counts/day normalized to 100 tons of target mass.

**5t 0.1% Gd-loaded  
scintillators**

**Shielding:**

**300 MWE**

**2 m scintillator +  
0.14m Fe**

**1km baseline**

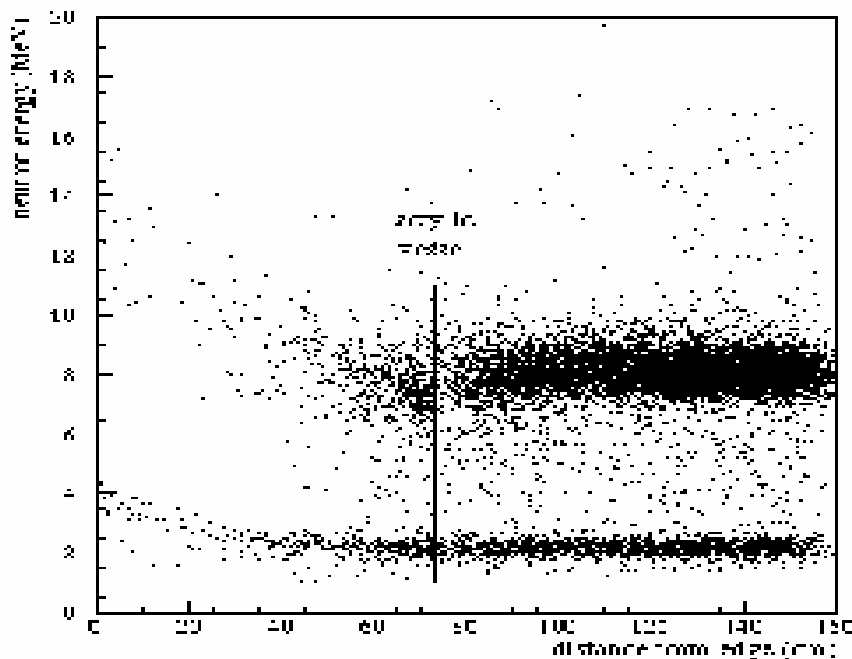
**Signal: ~30/day**

**Eff. : ~70%**

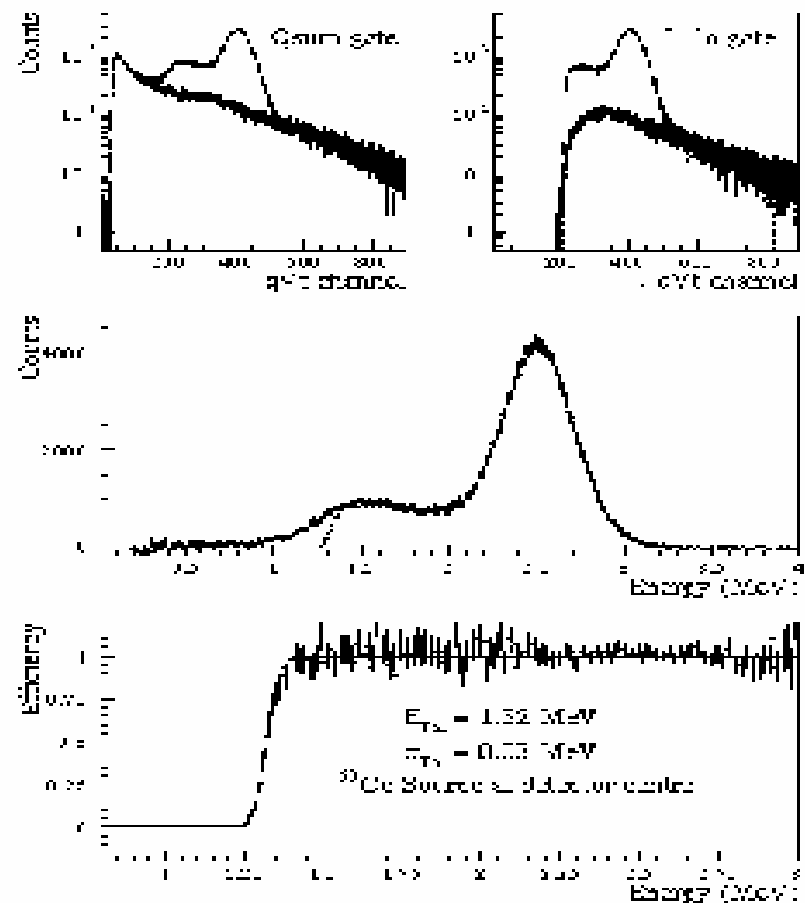
**BK:**

**corr. 1/day**

**uncorr. 0.5/day**



Position cut



Energy cut

# Systematics

<b>sources</b>	<b>Relative error (%)</b>
<b>Reaction cross section</b>	<b>1.9</b>
<b>Number of protons</b>	<b>0.8</b>
<b>Detection efficiency</b>	<b>1.5</b>
<b>Reactor power</b>	<b>0.7</b>
<b>Energy released per fission</b>	<b>0.6</b>
<b>total</b>	<b>2.7</b>

# Closer look -- Detection efficiency

Table 6. Summary of the neutrino detection efficiencies.

selection	$\epsilon(\%)$	rel. error (%)
positron energy*	97.8	0.8
positron-geode distance	99.9	0.1
neutron capture	84.6	1.0
capture energy containment	94.6	0.4
neutron-geode distance	99.5	0.1
neutron delay	93.7	0.4
positron-neutron distance	98.4	0.3
neutron multiplicity*	97.4	0.5
combined*	69.8	1.5

\* average values



# Experience gained

- Not stable Gd-loaded scintillator ( $\lambda \sim 5-2\text{m}$ )
- PMT directly in contact with scintillator  $\rightarrow$  too high uncorr. Background  $\rightarrow$  too high  $E_{\text{th}}$  (1.32 MeV)
- Good shielding  $\rightarrow$  low background
- Homogeneous detector  $\rightarrow$  Gd peak at 8 MeV
- 2m scintillator shielding gives a neutron reduction of  $0.8 \cdot 10^6$ .

**12t 0.1% Gd-loaded  
scintillators**

**Shielding:**

**32 MWE /1m water**

**0.9 km baseline**

**Signal:      ~20/day**

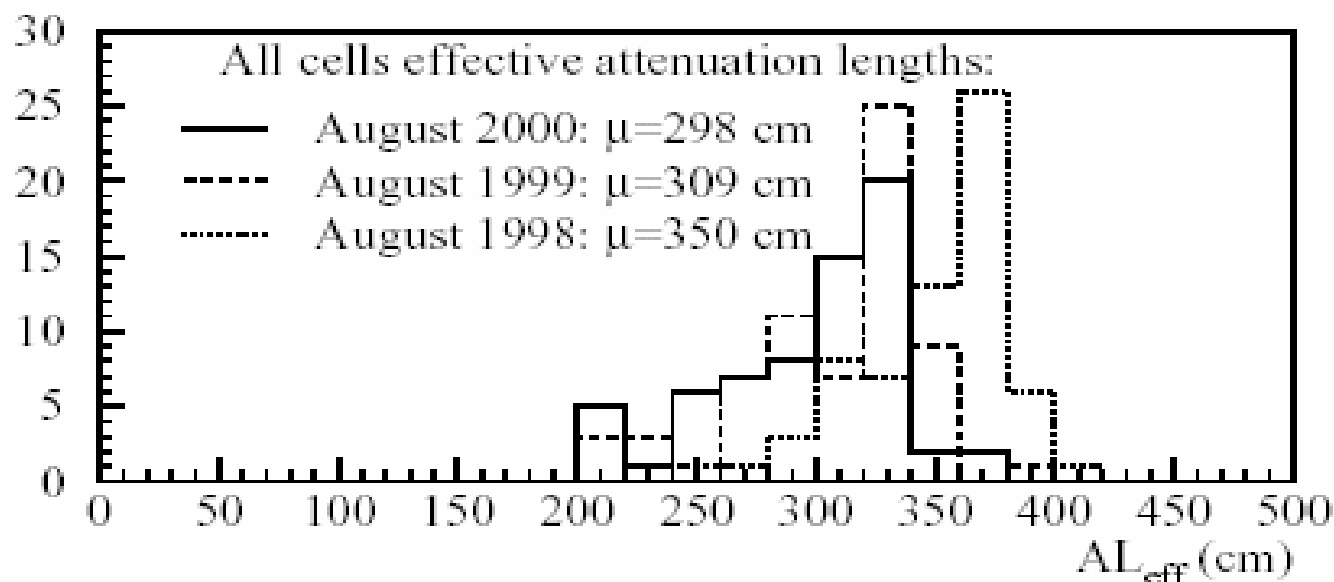
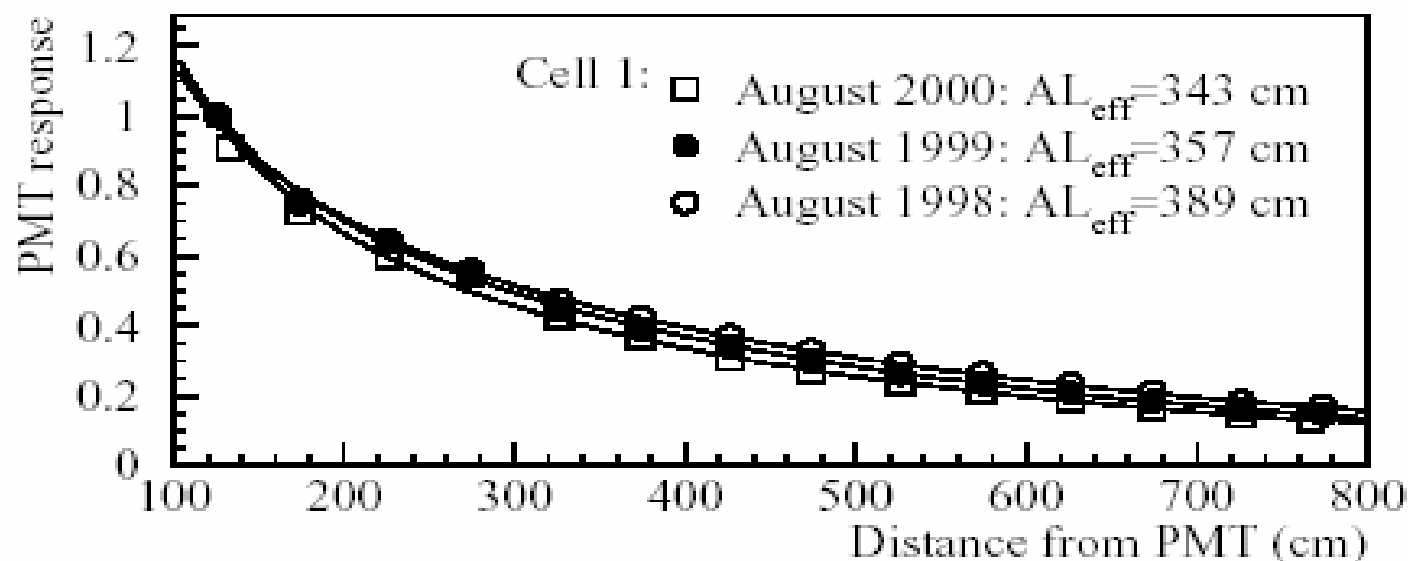
**Eff.            ~ 10%**

**BK:**

**corr.:        ~ 15/day**

**uncorr.      ~ 7/day**

# Very stable Gd-loaded liquid scintillator



# Systematics

Sources	power method	Swap method
e+ trigger efficiency	2.0	2.0
n trigger efficiency	2.1	2.1
$\nu$ flux prediction	2.1	2.1
$\nu$ selection cuts	4.5	2.1
Background variation	2.1	N/A
$(1-\varepsilon_1) B_{\text{pn}}$	N/A	3.3
Total	6.1	5.3

Error on  $\nu$  selection cuts obtained from multi-variable analysis

# Experience gained

- Good Gd-loaded scintillator( $\lambda \sim 11\text{m}$ )
- Not enough shielding  $\rightarrow$  too high corr./uncorr. Background
- Segmentation makes Gd capture peak  $< 6\text{MeV}$   $\rightarrow$  too high uncorr. Background
- Rn may enter the detector, problem ?
- Veto eff. is not high enough(97.5%)
- Swap method to measure/cancel backgrounds  $\rightarrow$  key to success
- 1m water shielding gives a neutron reduction of  $10^6$  (lower energy, complicated event pattern).

# KamLAND

1000t scintillators

Shielding:

3000 MWE/3m Water

180 km baseline

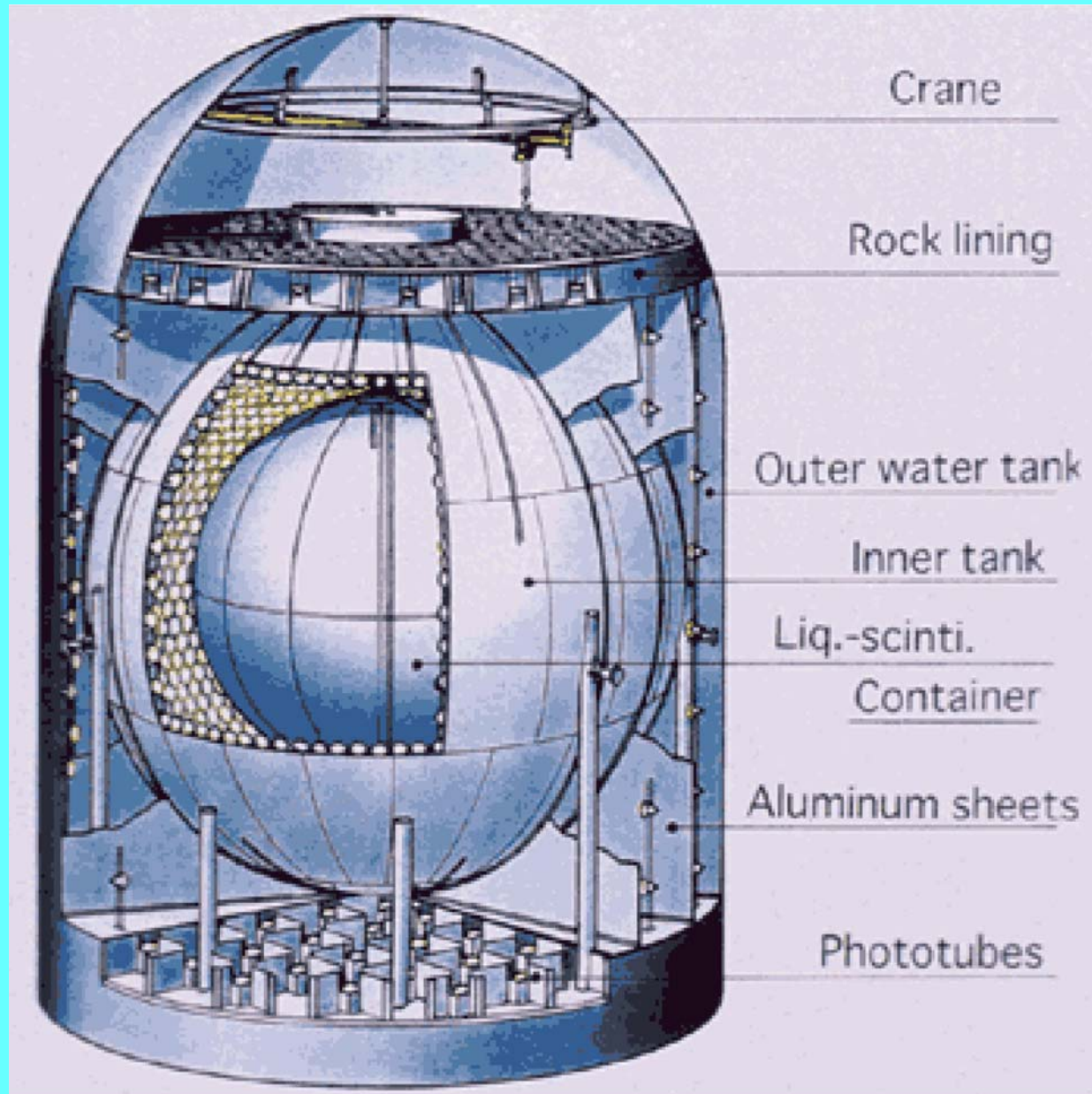
Signal:  $\sim 0.5/\text{day}$

Eff.  $\sim 40\%$

BK:

corr.:  $\sim 0.001/\text{day}$

uncorr.  $\sim 0.01/\text{day}$

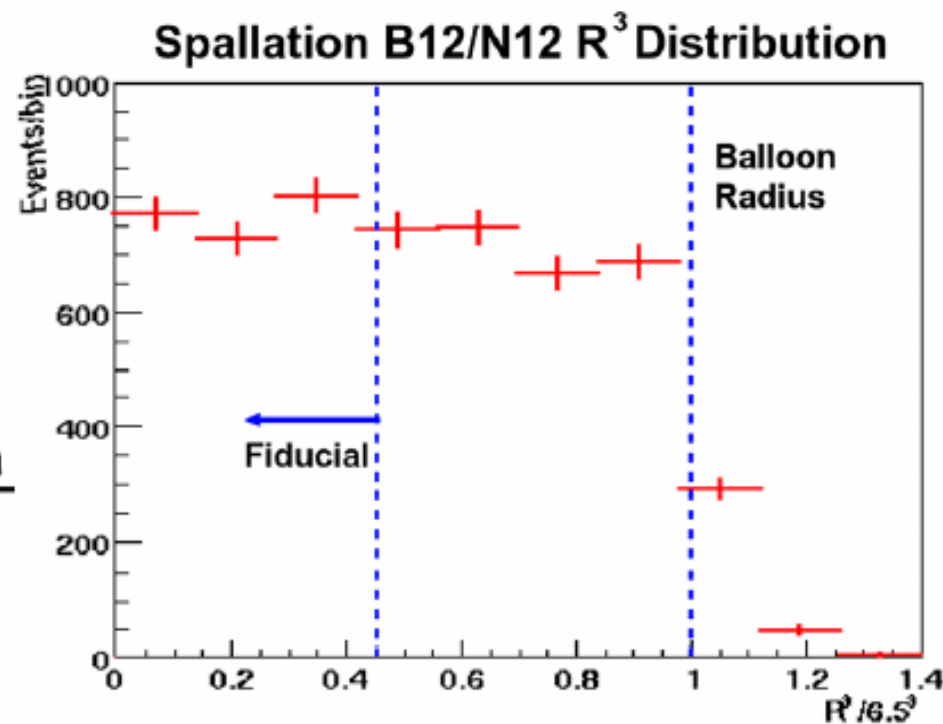
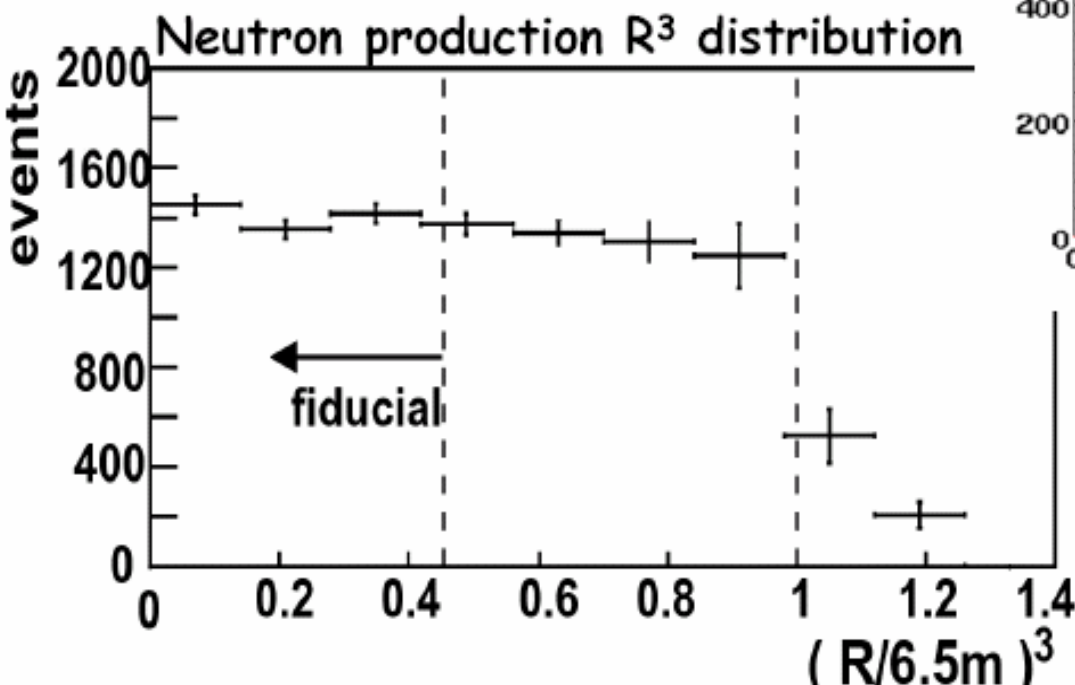


# Large error on fiducial volume

$$R_{fid} = 6.5 \text{ m}$$

$$R_{fid} = 5 \text{ m}$$

$$\Delta V_{fid}/V_{fid} = 4.6 \%$$



*Fiducial mass 408 ton  
(out of 1000)*



# Systematic errors

Systematic errors	E>0.9 MeV	E>2.6 MeV
Total LS mass	2.13	2.13
Fiducial mass ratio	4.03	4.03
Energy threshold	--	2.13
Efficiency of cuts	2.06	2.06
Live time	0.07	0.07
Reactor power	2.05	2.05
Fuel composition	1.0	1.0
Time lag	0.28	0.28
$\nu$ spectra	2.25	2.48
Cross section	0.2	0.2
Total	6.0	6.4

# Experience gained

- Very good shielding
- Balloon not good → target mass not well defined
- Light transport in scintillator unknown → particularly bad for Large detectors → large error on position reconstruction
- Background from  $^8\text{He}/^9\text{Li}$
- Not good enough veto tracking system
- 3m water shielding gives a neutron reduction of  $>13 \times 10^6$  (high energy).

# Three main types of errors:

reactor related(3%)

background related(0.5-4%)

detector related(3%)

How to reduce these errors ?

Can we do better than 1% ???

# Systematic error comparison

		Chooz	Palo Verde	KamLAND
Reactor power		0.7	0.7	2.05
Reactor fuel/v spectra		2.0	2.0	2.7
v cross section		0.3	0.2	0.2
No. of protons	H/C ratio	0.8	0.8	1.7
	Mass	-	-	2.1
Efficiency	Energy cuts	0.89	2.1	0.26
	Position cuts	0.32		3.5
	Time cuts	0.4		0.
	P/Gd ratio	1.0		-
	n multiplicity	0.5		-
background	correlated	0.3	3.3	1.8
	uncorrelated	0.3	1.8	0.1
Trigger		0	2.9	0
livetime		0	0.2	0.2

# Important point to have small systematic error

- Energy threshold less than 0.9 MeV
- Homogeneous detector
- Scintillator mass well determined
- Target scintillator all from one batch, mixing procedures well controlled
- Not too large detector
- Comprehensive calibration program
- Background well controlled → good shielding
- Be able to measure everything (Veto ineff., background, energy/position bias, ...)
- A lot of unforeseen effects will occur when looking at 0.1% level

# Reactor related error

- Use two detectors, far/near to cancel reactor related errors and some of detector/background related errors
- For one/two reactors, cancellation is exact
- For multiple reactors, cancellation is NOT exact
  - Multiple near detectors may be needed
  - Only uncorrelated errors contribute to final errors
  - Optimum positions to have minimum errors, typically 0.1-0.2%

➔ see J. Cao's talk

# Background related error

- Just to have enough shielding
- How much is enough ?
  - Uncorrelated backgrounds: U/Th/K MC
  - Correlated backgrounds:

Y.F. Wang et al., PRD64(2001)0013012

T. Hagner et al., Astroparticle Phys. 14(2000) 33



# Background - correlated

- Cosmic-muon-induced neutrons:
  - $B/S < 0.005 \rightarrow 1/\text{day} @ \sim 1\text{km}$
  - Can be measured by veto tagging, accuracy  $< 20\%$
  - Veto rate  $< 1\text{KHz}$ , 2-3 layers RPC(1600-2400 m<sup>2</sup>) ?
  - Methods:
    - Overburden  $> 100 \text{ MWE}$
    - Active Veto, ineff.  $< 0.5\%$ , known  $< 0.2\%$
  - Three scenarios:

	100 MWE	300 MWE	1000 MWE
muon rate/m <sup>2</sup> (Hz)	4	0.4	0.02
n rate in rock/m <sup>3</sup> (/day)	11000	1600	160
reduction required (10 <sup>6</sup> )	9.2	1.4	0.14
Shielding (water equivalent) (m)	2.5m	2.1m	1.5m

# Other correlated backgrounds

- $\beta$ -neutron instable isotopes from cosmic  $\mu$ 
  - $^8\text{He}/^9\text{Li}$ ,  $\text{Br}(n) = 12\%/48\%$ ,  $^9\text{Li}$  dominant
  - Production rates =  $f_\mu \cdot N_A \cdot \sigma \cdot \text{Br}$

	100 MWE	300 MWE	1000 MWE
Average $E_\mu$ (GeV)	36	64	160
muon rate/m <sup>2</sup> (Hz)	4	0.4	0.02
Cross section ( $\mu\text{b}$ )	0.61	0.94	1.86
$^8\text{He}/^9\text{Li}$ (1/day/module)	3.4	0.53	0.053

- Depth > 300 MWE, best 1000 MWE

# Background - Uncorrelated

- $B/S < 0.05 \rightarrow < 8/\text{day}$  @ far site
- Can be measured by swap method, precision  $\sim \sqrt{B/s}=2.5\%/\text{day}$
- single rate @  $0.9\text{MeV} < 50\text{Hz}$ 
  - $2 \cdot R_\gamma \cdot R_n \cdot \tau < 0.04/\text{day/module}$
- Methods:
  - Low activity glass for PMT,  $> 0.5\text{m}$  oil shielding (dominant!)
  - 3 MWE shielding, low activity sand/aggregate or Fe ?
  - Rn concentration  $< 20 \text{ Bq/m}^3$ ,  $\text{N}_2$  flushing ?
  - (U, Th, K) in Scintillator  $< 10^{-13} \text{ g/g}$ , clean Gd
  - All mechanical structure made of low activity materials
  - Calibration gadget made of clean materials such as Teflon, ...
  - Clean everywhere, no dust, no ...

# Systematic error comparison

		Chooz	Palo Verde	KamLAND	Cancel ?
Reactor power		0.7	0.7	2.05	<0.1%
Reactor fuel/v spectra		2.0	2.0	2.7	
v cross section		0.3	0.2	0.2	0
No. of protons	H/C ratio	0.8	0.8	1.7	0
	Mass	-	-	2.1	<0.1
Efficiency	Energy cuts	0.89	2.1	0.26	0.2
	Position cuts	0.32		3.5	0.2
	Time cuts	0.4		0.	0.1
	P/Gd ratio	1.0		-	0
	n multiplicity	0.5		-	<0.1
background	correlated	0.3	3.3	1.8	<0.1
	uncorrelated	0.3	1.8	0.1	<0.1
Trigger		0	2.9	0	<0.1
livetime		0	0.2	0.2	<0.1

# Possibly best systematic errors

- Reactor  $< 0.1\%$
- Background  $< 0.2\%$
- Energy cut  $\sim 0.2\%$
- Position cut  $\sim 0.2\%$
- Time cut  $< 0.1\%$
- Livetime  $\sim 0.1\%$
- Other unexpected  $< 0.2\%$
- Total  $< 0.5\%$

# Further reduction of systematic errors: multiple modules

- **Smaller modules have less unknowns**
- **Multiple handling to control systematic errors**
- **Easy construction**
- **Easy movable detector**
- **Scalable**
- **Easy to correct mistakes**

# Summary

- **Systematic errors from reactor well under control:  
Near vs Far**
- **Errors from backgrounds: just need “enough”  
shielding**
- **Errors from detector: can be controlled to 0.5%  
level if the detector carefully designed**

**$\theta_{13}$  experiment at 0.5% level is possible !**

**Let's do it !!!**