

θ_{13} Quest at Daya Bay

Measurement of θ_{13} Mixing Parameter

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on behalf of Daya Bay Experiment Collaboration



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Nu HoRlizons III, February 8-10, 2010, Allahabad, India



- 1 Motivation
- 2 Location and Onsite Layout
- 3 Detection Method
- 4 Systematic Uncertainties
- 5 Backgrounds
- 6 Sensitivity
- 7 Status and Plans

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PMNS Matrix

$$\underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix}}_{\theta_{23} \approx 45^\circ} \times \underbrace{\begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix}}_{\theta_{13} < 10^\circ} \times \underbrace{\begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\theta_{12} \approx 35^\circ}$$

$\theta_{23} \approx 45^\circ$

Atmospheric ν

Accelerator ν

$\theta_{13} < 10^\circ$

Short-baseline Reactor ν

Future accelerator ν

$\theta_{12} \approx 35^\circ$

Solar ν

Long-baseline Reactor ν

- Measure θ_{13} with sensitivity of $\sin^2 2\theta_{13} < 0.01^1$ at 90% C.L.
- Currently known to be $\sin^2 2\theta_{13} < 0.19$ at 90% C.L. from the Chooz Reactor Neutrino Experiment in France

¹ $\sin^2 \theta_{13} < 0.0025$ and $\theta_{13} < 3^\circ$

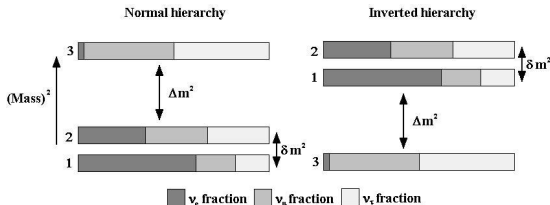
Importance

- Step to complete basic model of ν oscillations
- Open gate to go further

- Is it possible to measure CP violation from neutrino oscillations:

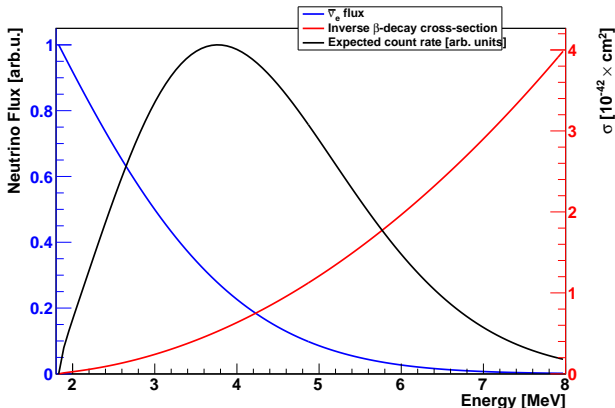
$$P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = \sin(2\theta_{12}) \sin(2\theta_{23}) \cos^2(\theta_{13}) \sin(2\theta_{13}) \sin \delta$$

- Mass hierarchy: $m_2 < m_3$ or $m_2 > m_3$

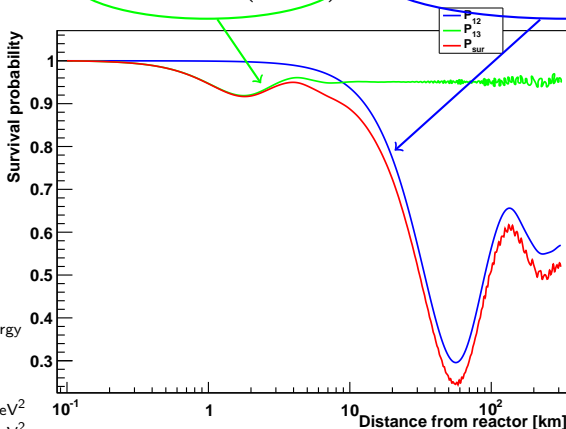


- Help discriminate among theoretical models of mixing matrix

$$P_{ee} \approx 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{13}^2 L}{4E_\nu} \right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left(\frac{\Delta m_{21}^2 L}{4E_\nu} \right)$$



$$P_{ee} \approx 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{13}^2 L}{4E_\nu} \right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left(\frac{\Delta m_{21}^2 L}{4E_\nu} \right)$$



Integrated over $\bar{\nu}_e$ energy

$E_\nu \in [1.8, 8]$ MeV

$\sin^2 2\theta_{13} = 0.1$

$\theta_{12} = 34^\circ$

$\Delta m_{13}^2 = 2.5 \times 10^{-3} \text{ eV}^2$

$\Delta m_{21}^2 = 7.9 \times 10^{-5} \text{ eV}^2$

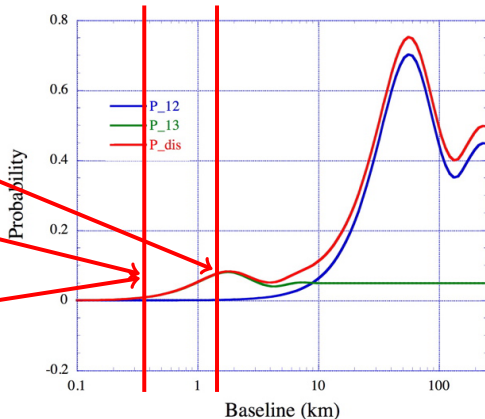
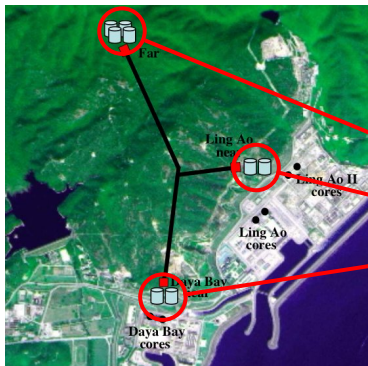
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Daya Bay Nuclear Powerplant
~ 50 km from Hong Kong



- 3 sites at different distances — 2x Near (2x2 detectors), 1xFar (4 detectors)



Distance from
detectors to
reactor cores in
meters

Reactors	Experimental site		
	DyB	LA	Far
DayaBay	363	1348	1986
LingAo I	857	481	1618
LingAo II	1307	526	1613
Overburden	98	112	355

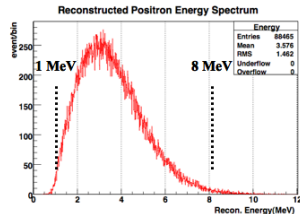
Reactor Thermal Output:
11.6 GW now, 17.4 GW in
2011

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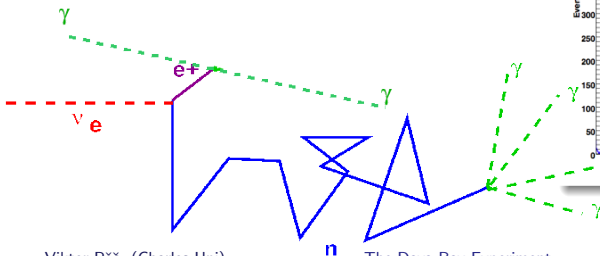
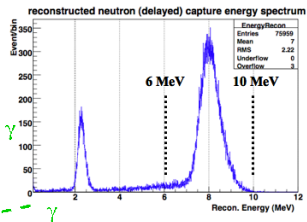
- Inverse β -decay :

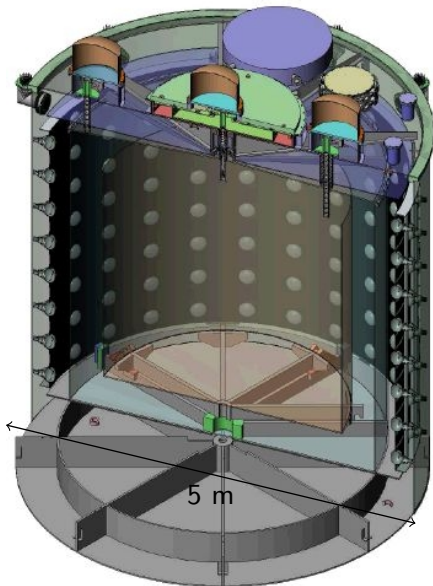
$$\bar{\nu}_e + p \rightarrow e^+ + n$$
- Trigger on 2-fold coincidence:
 - Prompt signal from e^+
 - Delayed signal from n capture on Gadolinium $\approx 30\mu s$
- Detector with Gd doped Liquid Scintillator (LS)

Prompt signal

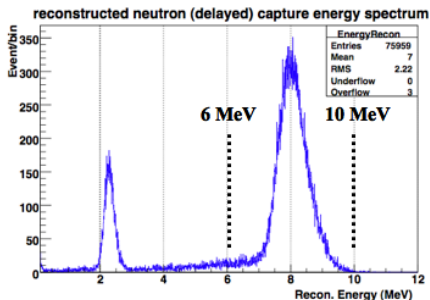
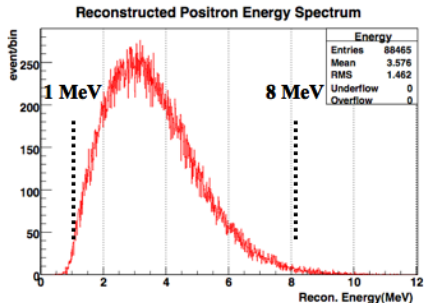


Delayed signal





- Cylindrical **3-Zone Structure** separated by acrylic vessels
 - **Target:** Inner 20t GdLS (0.1% of Gd, $d=3\text{m}$)
 - **γ -catcher:** Mid 20t LS ($d=4\text{m}$, $\approx 42\text{cm}$ thick)
 - **Oil Buffer:** Outer 40t mineral oil ($d=5\text{m}$, $\approx 49\text{cm}$ thick)
- 192 8-inch PMTs
- $12\%/\sqrt{E(\text{MeV})}$ energy resolution
- Reflectors on top and bottom



Site	Signal/day/module
Daya Bay	840
Ling Ao	760
Far Site	90

- Positron energy cuts at 1 – 8 MeV
- Neutron capture energy cut at 6 MeV
- Time cut 0.3 – 200 μ s

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$$\frac{N_f}{N_n}(E) = \frac{N_{p,f}}{N_{p,n}} \left(\frac{L_n}{L_f} \right)^2 \frac{\epsilon_f}{\epsilon_n} \frac{P(E, L_f)}{P(E, L_n)}$$

- Expected ratio of measured events for particular energy at Near and Far site
- Number of protons in target – careful measurements during filling
- Neutrino flux at distance $L \propto 1/L^2$
- Detector detection efficiencies – intensive calibration program
- Survival probabilities – sign of oscillation

Detector
related
uncertainty

Source of uncertainty		Detector Systematic Uncertainties	
		Conservative	Goal
# protons		0.3	0.1
Detector Efficiency	Energy cuts	0.2	0.1
	Time cuts	0.1	0.03
	H/Gd ratio	0.1	0.1
	n multiplicity	0.05	0.05
	Trigger	0.01	0.01
	Live time	<0.01	<0.01
Total		0.38%	0.18%

Reactor
related
uncertainty

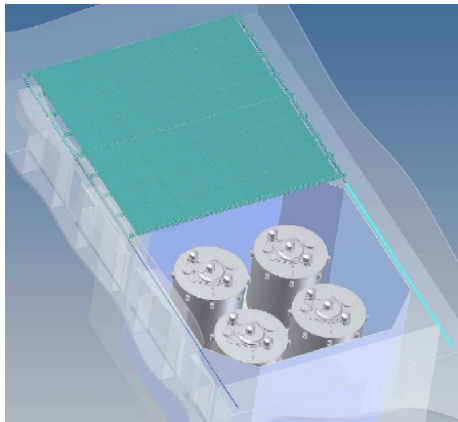
Number of cores	Power	Location	Total
4	0.035%	0.08%	0.087%
6	0.097%	0.08%	0.126%

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- **Accidental** coincidence of uncorrelated signals — natural radioactivity
- Correlated signals from **fast neutrons** — spallation processes of muons in surrounding rock
- β -delayed neutron decays of ${}^9\text{Li}$ and ${}^8\text{He}$ — products of muonic showers

Background subtraction uncertainties

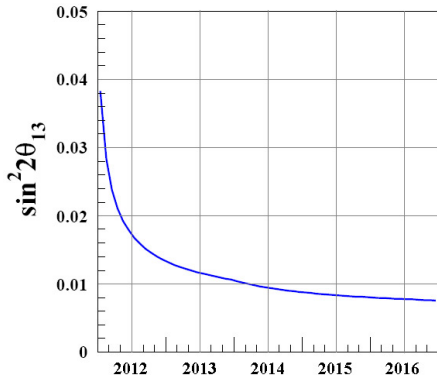
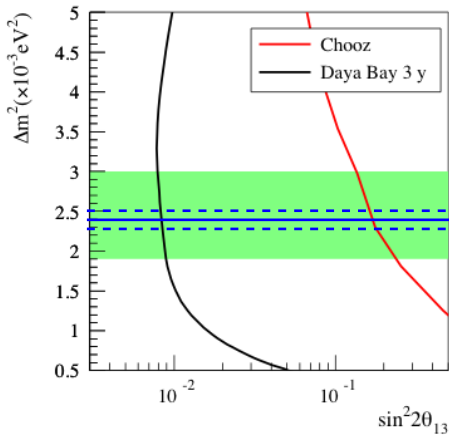
	DYB	LA	Far
Fast n/signal	0.1%	0.1%	0.1%
${}^9\text{Li}$, ${}^8\text{He}$ /signal	0.3%	0.2%	0.2%
Accidentals/signal	<0.2%	<0.2%	<0.1%



- Multiple muon veto detectors
- **Water Čerenkov**
 - ADs submerged in water, provide $\geq 2.5\text{m}$ shielding against radioactivity
 - **Inner/Outer** regions optically separated
 - 8-inch PMTs on frames (289/near, 384/far site)
- **RPC—Resistive Plate Chamber**
 - 4 layers in modules
 - Layer of modules covers water pool
 - Provides independent veto system
- Combined efficiency of both systems $> 99.5\%$

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- 90% C.L. after 3 years of data taking assuming baseline systematics, compared to Chooz results



$$\Delta m^2 = 0.0025 \text{ eV}^2$$

- Green band is 90% C.L. limits from atmospheric neutrino experiments
- Best fit and 1σ errors from MINOS
ref. arXiv:0808.2016v2

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Excavation continues, more than 3,000 m of tunnels excavated



2 experimental halls excavated



First AD being assembled



- October 2007: Ground breaking
- March 2009: Surface assembly building occupancy
- Upcoming months: Commissioning first AD - Dry run
- 2010: Daya Bay Near Hall ready for data
- 2011: Far Hall ready for data



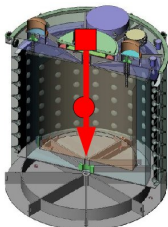
Region	Institutions	Members
China	14	109
Czech	1	4
Hong Kong	2	16
Russia	2	5
Taiwan	3	13
USA	16	96
Sum	38	243

Purpose

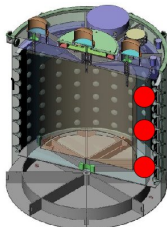
- PMT readout and system integration
- Coordination of calibration and DAQ
- PMT measurements
 - Timing
 - Gain
 - Peak-to-valley ratio
 - Relative efficiency

Implementation

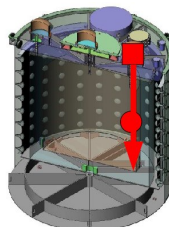
- Assembled detector without filling (no scintillator, no mineral oil)
- Use LED sources



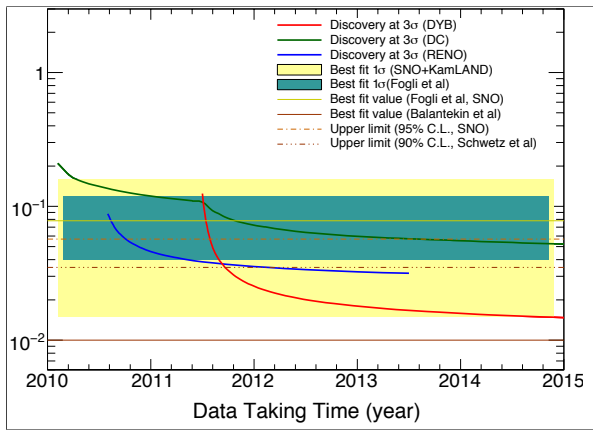
Viktor Pěč (Charles Uni)



The Daya Bay Experiment



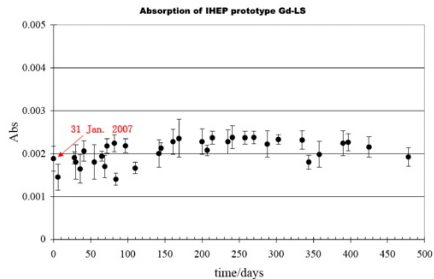
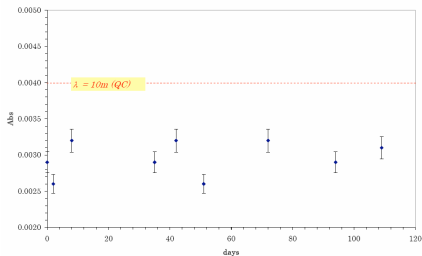
Nu HoRlizons III



- Analysis by C. Lewis, B. Littlejohn, M. McFarlane, W. Wang, K. Heeger from University of Wisconsin
- Models for RENO and Double Chooz from GLOBES website
- Starting dates from Huber et al (arXiv: 0907.1896)

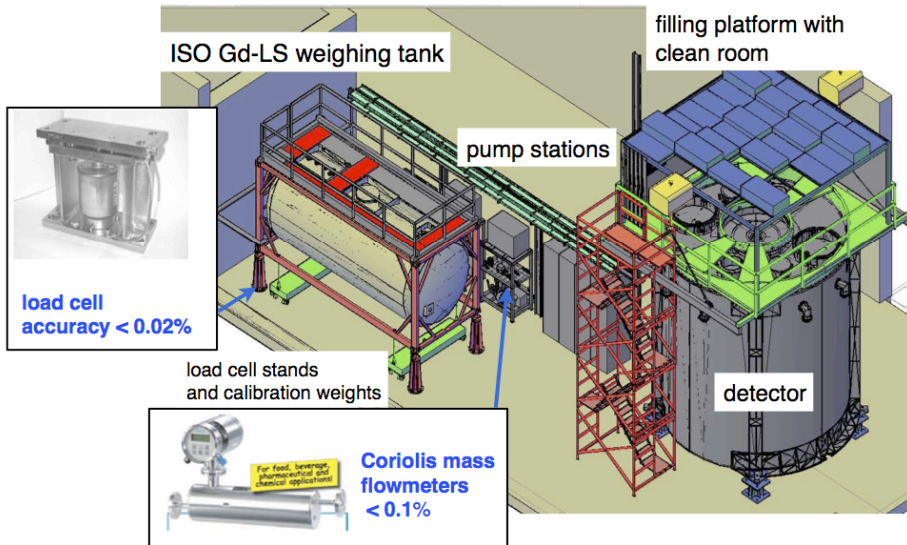
	DYB site	LA site	Far site
Vertical overburden (m)	98	112	355
Vertical overburden (m.w.e.)	255	291	910
Muon Flux (Hz/m ²)	1.16	0.73	0.041
Muon Mean Energy (GeV)	55	60	138

Source of uncertainty		Chooz (absolute)	Daya Bay (relative)		
			Conservative	Goal	Goal w/Swapping
# protons		0.8	0.3	0.1	0.006
Detector Efficiency	Energy cuts	0.8	0.2	0.1	0.1
	Position cuts	0.32	0.0	0.0	0.0
	Time cuts	0.4	0.1	0.03	0.03
	H/Gd ratio	1.0	0.1	0.1	0.0
	n multiplicity	0.5	0.05	0.05	0.05
	Trigger	0	0.01	0.01	0.01
	Live time	0	<0.01	<0.01	<0.01
Total uncertainty (detector-related)		1.7%	0.38%	0.18%	0.12%



- Weekly deployment of radioactive sources
 - β^+ – ^{68}Ge
 - neutrons – Am/Pu-C
 - γ – ^{60}Co , ^{137}Cs
- LED diffuser balls
 - monitor optical properties of materials
 - PMT gains and timing
 - Electronics performance

AD Filling and Target Mass Measurement

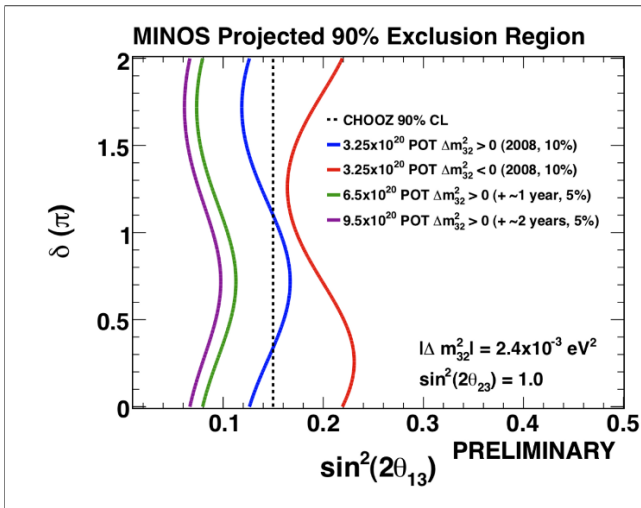


Huber et al., arxiv 0907.1896

Setup	t_ν [yr]	$t_{\bar{\nu}}$ [yr]	P_{Th} or P_{Target}	L [km]	Detector technology	m_{Det}
Double Chooz	-	3	8.6 GW	1.05	Liquid scintillator	8.3 t
Daya Bay	-	3	17.4 GW	1.7	Liquid scintillator	80 t
RENO	-	3	16.4 GW	1.4	Liquid scintillator	15.4 t
T2K	5	-	0.75 MW	295	Water Cerenkov	22.5 kt
NO ν A	3	3	0.7 MW	810	TASD	15 kt

Table 1: Summary of the standard setups at their nominal luminosities.

MINOS ν_e appearance observation



From MINOS web page.