



Measuring $\sin^2 2\theta_{13}$ with the Daya Bay nuclear power reactors

Yifang Wang

for Daya Bay collaboration

Institute of High Energy Physics

Daya Bay Neutrino Experiment

- Measure $\sin^2 2\theta_{13}$ with a sensitivity of 0.01 at 90% CL, an improvement of an order of magnitude over previous experiments
- 4 reactor cores, 2 more in 2011, a total of 17.4 GW
- Mountains near by, easy to construct a lab with enough overburden to shield cosmic-ray backgrounds
- 60 km from Hong Kong, Convenient Transportation, Living conditions, communications



Layout of the experiment



Near-Far detector schemes:

To cancel reactor-related errors

Residual error ~0.1%

Swap near-far detectors

To cancel detector-related errors.

Residual error ~0.2%

Detector deep undergrounds

To reduce backgrounds

B/S at near site: ~0.5%

B/S at far site: ~0.2%

Fast Measurement

DYB+Mid, 2008-2009

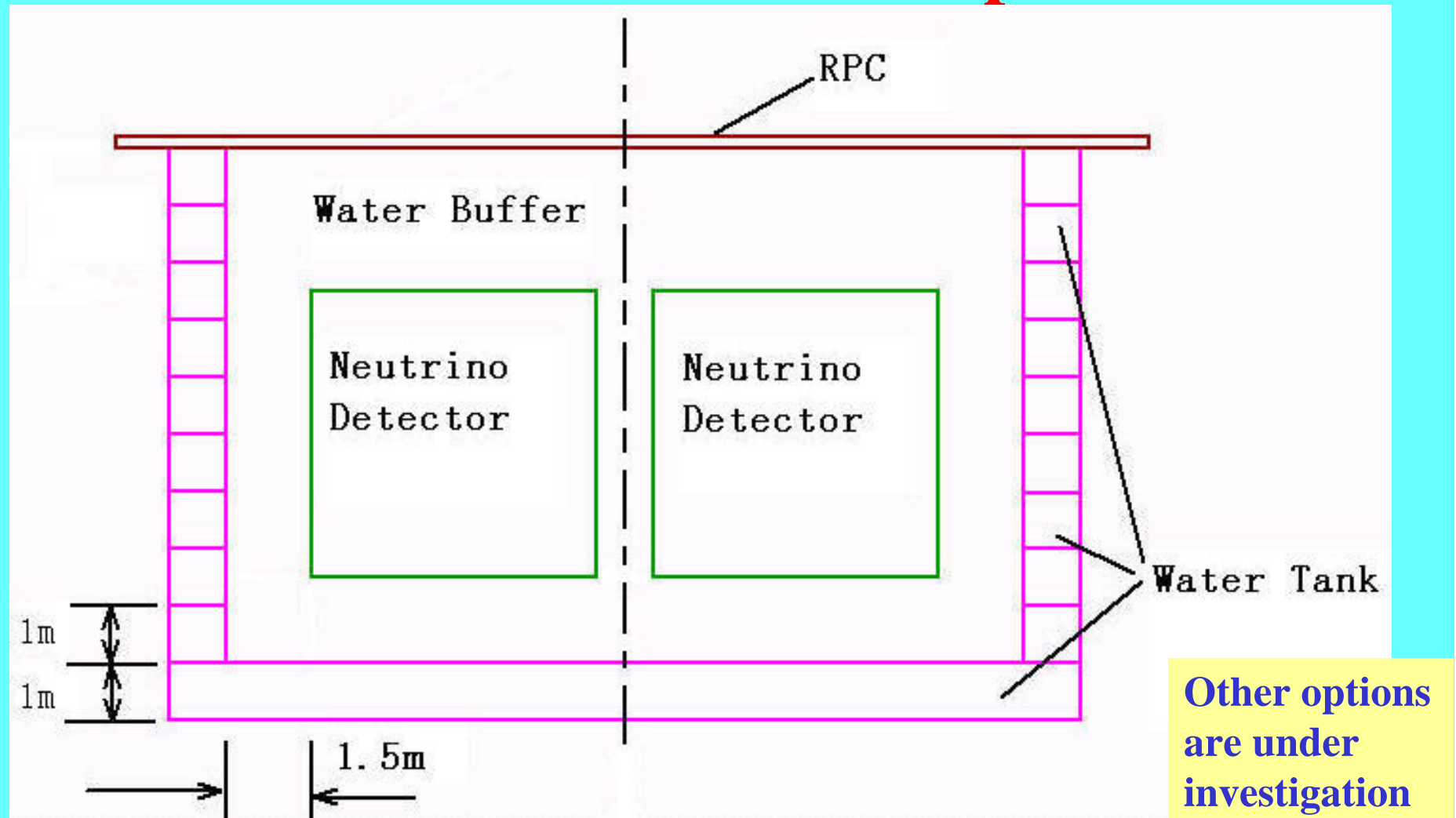
Sensitivity (1 year) ~0.03

Full Measurement

DYB+LA+Far, from 2010

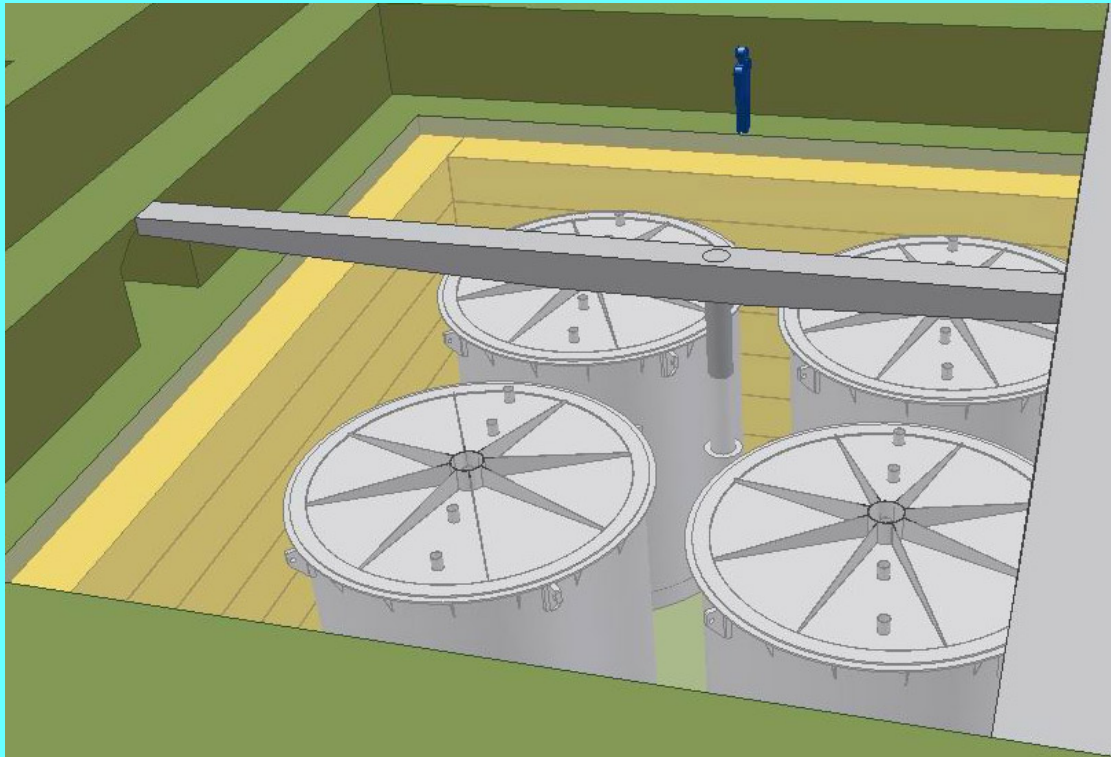
Sensitivity (3 year) <0.01

Baseline detector design: multiple neutrino modules and multiple vetos



Redundancy is a key for the success of this experiment

Neutrino detector: multiple modules

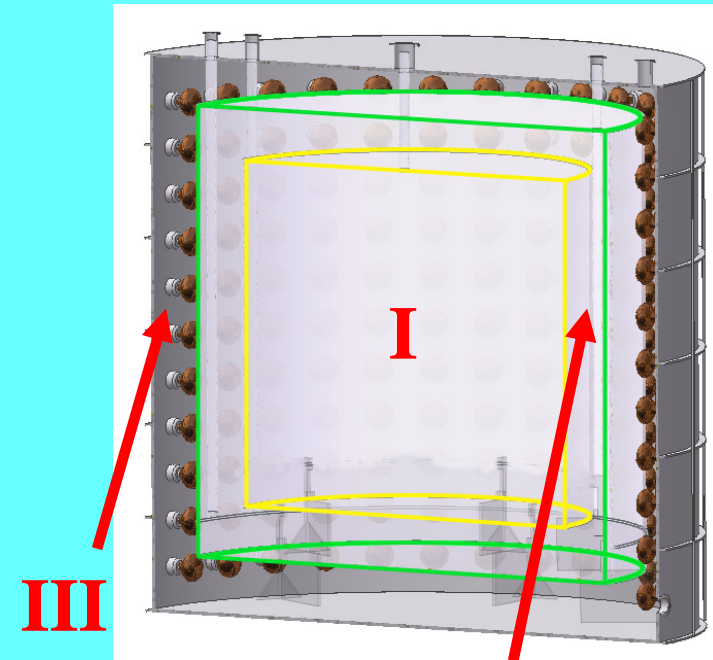


Two modules at
near sites
Four modules at
far site

- Multiple modules for side-by-side cross check
- Reduce uncorrelated errors
- Smaller modules for easy construction, moving, handing, ...
- Small modules for less sensitivity to scintillator aging, details of the light transport, ...

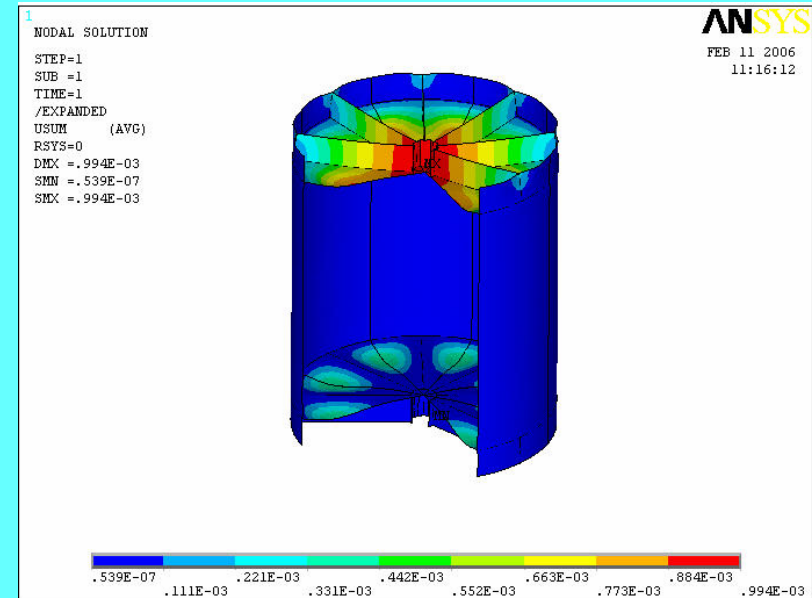
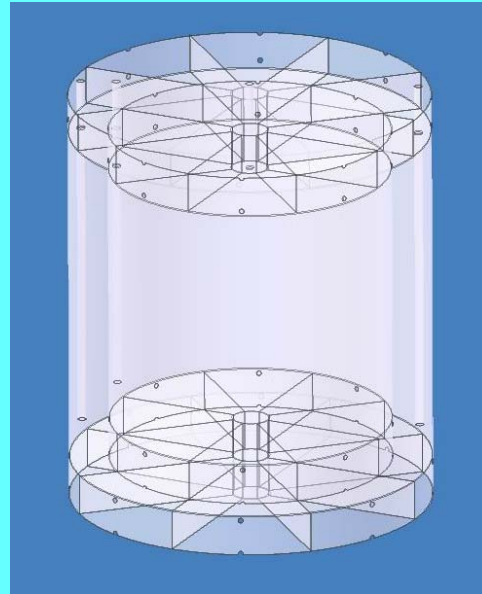
Central Detector modules

- **Three zones modular structure:**
 - I. **target: Gd-loaded scintillator,**
 - II. **γ -ray catcher: normal scintillator**
 - III. **Buffer shielding: oil**
- **Advantages: neutrino events on target are determined by capture time, not position**
- **Cylindrical module for easy construction**
- **Light reflector at top and bottom for cost saving**
- **Module dimension:**
 - **Target: 3.2 m high, 3.2 m diameter,**
determined by the limit of statistical errors
 - **γ -ray catcher: 0.45 m thick, determined by the limit of efficiency error**
 - **Buffer: 0.45 cm thick, determined by backgrounds from PMT glass**
- **~ 200 8" PMT/module**
- **Photocathode coverage: 5.6 % \rightarrow 10%(with light reflector)**
- **Performance: energy resolution 5% @ 8 MeV, position resolution ~ 14 cm**



Acrylic tanks

Design:

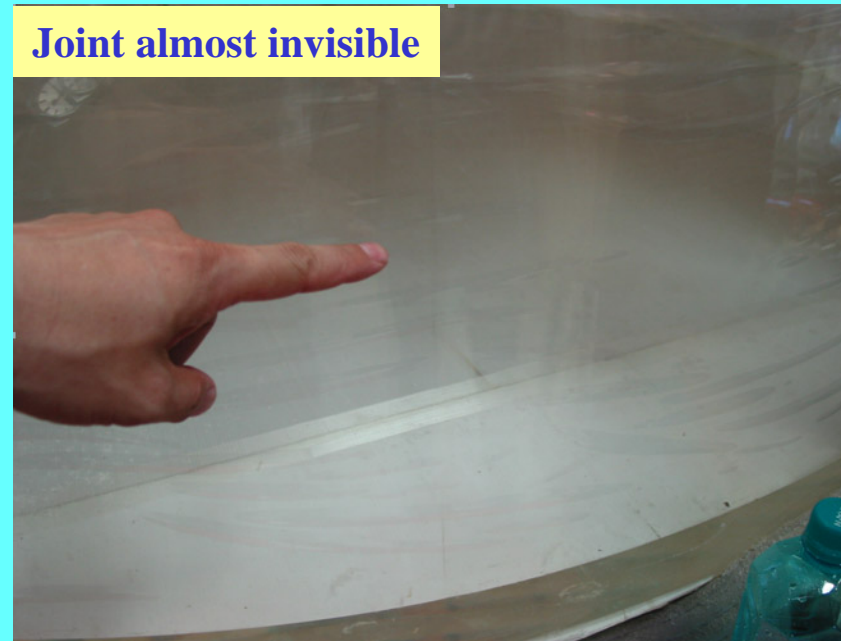


Manufacture

A 2m cylinder



Joint almost invisible



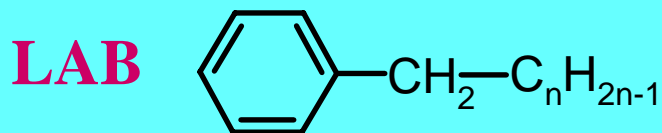
R&D of Gd-loaded scintillator

Requirements:

- Gd-loading: ~ (0.1-0.15)%
- light yield: > 50% Anthracene
- Attenuation length: > 10m
- Stability: < 3%/year

R&D Efforts:

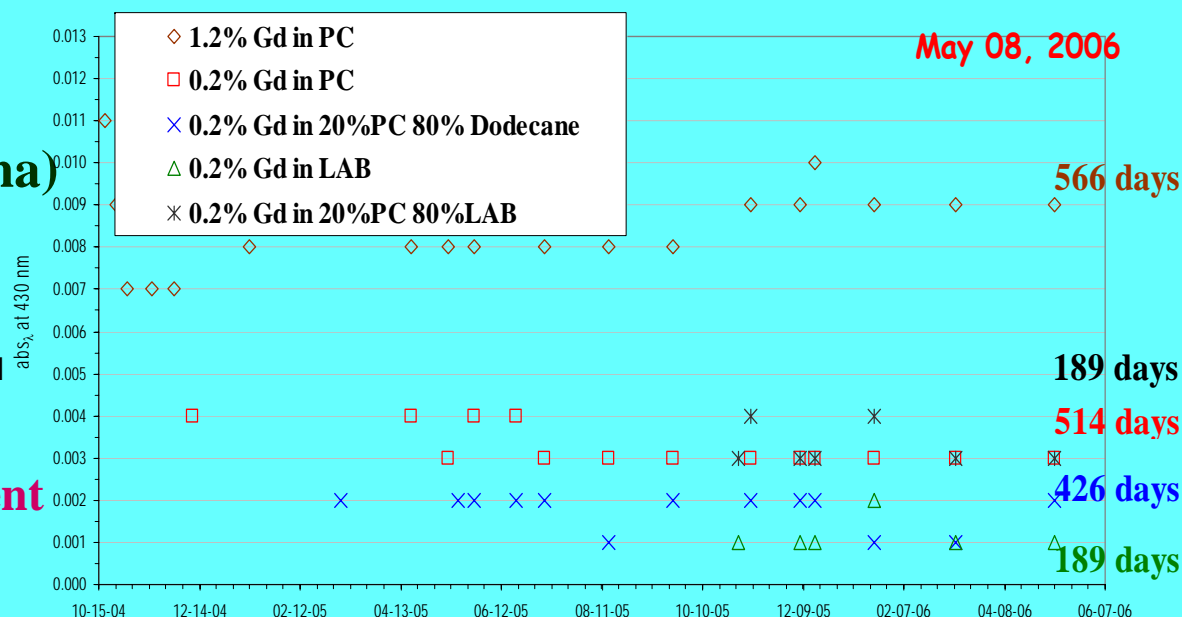
- PC+MO/DC(BNL)
- MT+MO/DC(IHEP)
- LAB+PC(BNL)
- LAB(IHEP, BNL, Dubna)



High light yield, very transparent
High flash point 147°C,
environmentally friendly
Low cost

Gd-loaded scintillator developed at IHEP

Sample	Atten. Len. (m)	Light yield
2: 8 mesitylene: dodecane(LS)	15.0	-
LAB	-	-
PPO+ bis-MSB(Flour) in LS	11.3	0.459
0.2% Gd-EHA + Flour + LS	8.3	0.528
Flour + LAB	23.7	0.542
0.2%Gd-TMHA+Flour+LAB	19.1	0.478



Stability of Gd-loaded scintillator developed at BNL

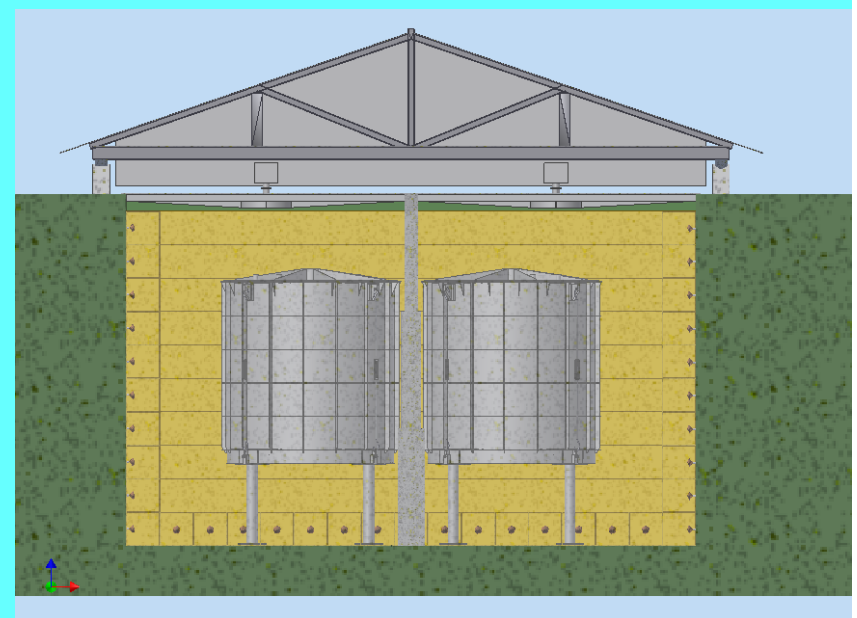
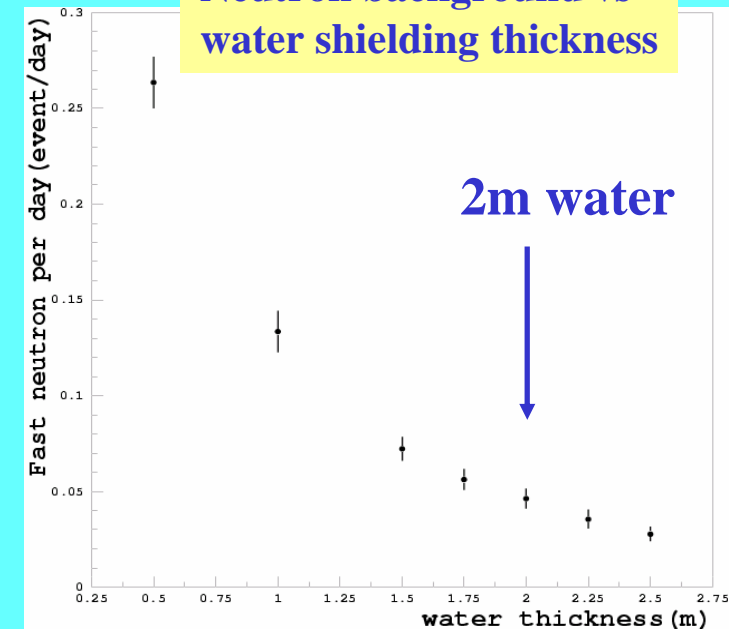
Calibration and Monitoring

- **Source calibration: energy scale, resolutions, ...**
 - **Deployment system**
 - **Automatic: quick but limited space points**
 - **Manual: slow but everywhere**
 - **Choices of sources: energy(0.5-8 MeV), activity(<1KHz), γ/n ,...**
 - **Cleanness**
- **Calibration with physics events:**
 - **Neutron capture**
 - **Cosmic-rays**
- **LED calibration: PMT gain, liquid transparency, ...**
- **Environmental monitoring: temp., voltage, radon, ...**
- **Mass calibration and high precision flow meters**
- **Material certification**

Water Buffer & VETO

- At least 2m water buffer to shield backgrounds from neutrons and γ 's from lab walls
- Cosmic-muon VETO Requirement:
 - Inefficiency < 0.5%
 - known to <0.25%
- Solution: Two active vetos
 - Active water buffer, Eff.>95%
 - Muon tracker, Eff. > 90%
 - RPC
 - Water tanks
 - scintillator strips
 - **total ineff. = 10%*5% = 0.5%**
- Two vetos to cross check each other and control uncertainty
- Baseline options:
Water pool + tracker

Neutron background vs water shielding thickness



Background related error

- Need enough shielding and active vetos
- How much is enough ? → **error < 0.2%**
 - Uncorrelated backgrounds: U/Th/K/Rn/neutron

Single gamma rate @ 0.9MeV < 50Hz

Single neutron rate < 1000/day

2m water + 50 cm oil shielding

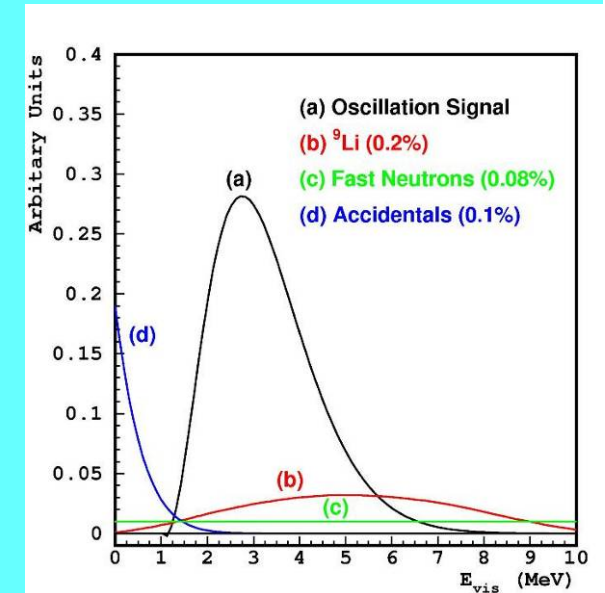
- Correlated backgrounds: $n \propto E_{\mu}^{0.75}$

Neutrons: >100 MWE + 2m water

Y.F. Wang et al., PRD64(2001)0013012

$^8\text{He}/^9\text{Li}$: > 250 MWE(near), >1000 MWE(far)

T. Hagner et al., Astroparticle. Phys. 14(2000) 33



	Near	far
Neutrino signal rate(1/day)	560	80
Natural backgrounds(Hz)	45.3	45.3
Accidental BK/signal	0.04%	0.02%
Correlated fast neutron Bk/signal	0.14%	0.08%
$^8\text{He}+^9\text{Li}$ BK/signal	0.5%	0.2%

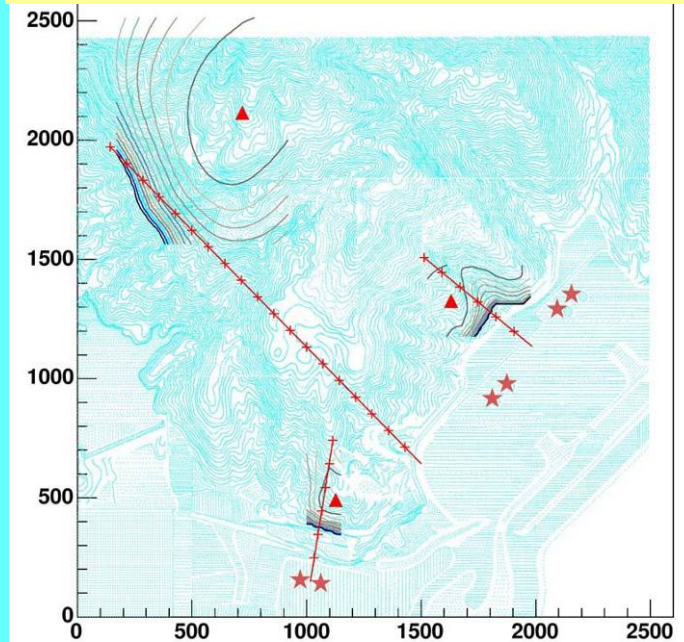
Systematic errors

		Chooz	Palo Verde	KamLAND	Daya Bay
Reactor power		0.7	0.7	2.05	<0.13%
Reactor fuel/v spectra		2.0	2.0	2.7	
v cross section		0.3	0.2	0.2	0
No. of protons	H/C ratio	0.8	0.8	1.7	0.2 → 0
	Mass	-	-	2.1	0.2 → 0
Efficiency	Energy cuts	0.89	2.1	0.26	0.2
	Position cuts	0.32		3.5	0
	Time cuts	0.4		0.	0.1
	P/Gd ratio	1.0		-	0.1 → 0
	n multiplicity	0.5		-	<0.1
background	correlated	0.3	3.3	1.8	0.2
	uncorrelated	0.3	1.8	0.1	<0.1
Trigger		0	2.9	0	<0.1
livetime		0	0.2	0.2	0.03

Baseline optimization and Sensitivity to $\sin^2 2\theta_{13}$

- Reactor-related correlated error: $\sigma_c \sim 2\%$
- Reactor-related uncorrelated error: $\sigma_r \sim 1\text{-}2\%$
- Neutrino spectrum shape error: $\sigma_{\text{shape}} \sim 2\%$
- Detector-related correlated error: $\sigma_D \sim 1\text{-}2\%$
- Detector-related uncorrelated error: $\sigma_d \sim 0.5\%$
- Background-related error:
 - fast neutrons: $\sigma_f \sim 100\%$,
 - accidentals: $\sigma_n \sim 100\%$,
 - isotopes(^8Li , ^9He , ...) : $\sigma_s \sim 50\text{-}60\%$
- Bin-to-bin error: $\sigma_{\text{b2b}} \sim 0.5\%$

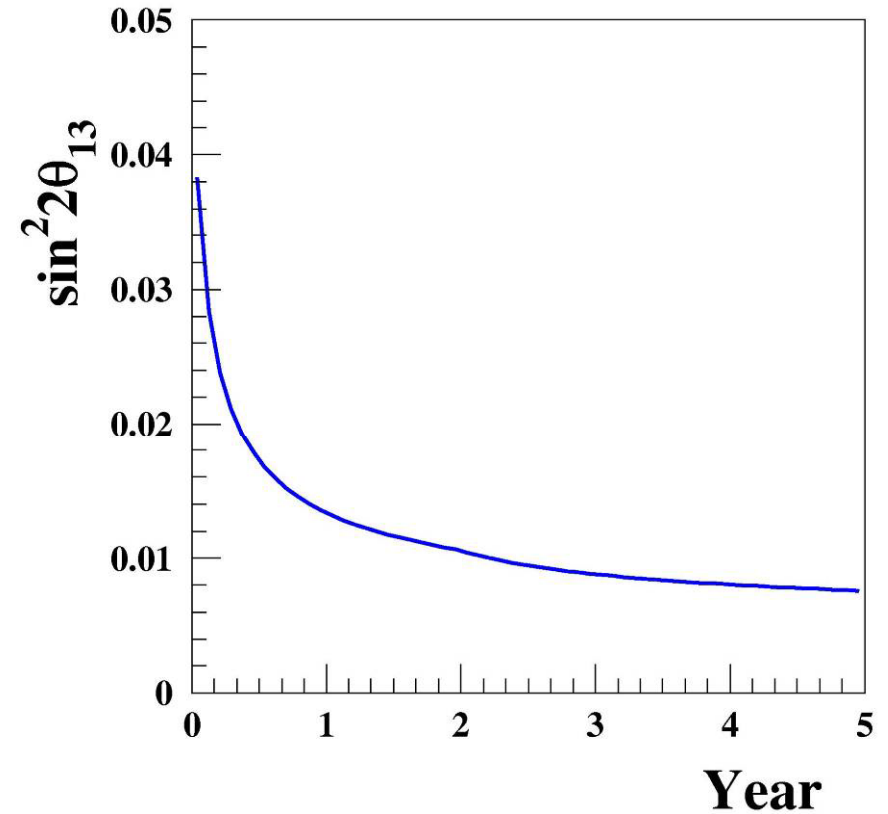
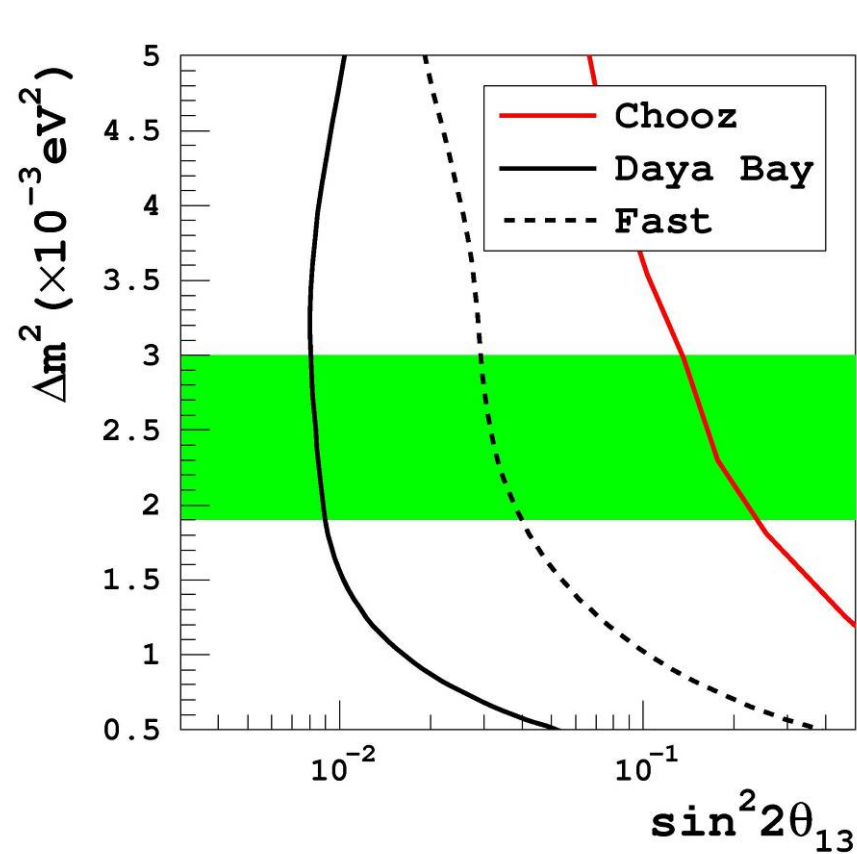
Many are cancelled by the near-far scheme and detector swapping



$$\chi^2 = \min_{\alpha's} \sum_{i=1}^{Nbin} \sum_{A=1,3} \frac{\left[M_i^A - T_i^A (1 + \alpha_D + \alpha_c + \alpha_d^A + c_i + \sum_r \frac{T_i^{rA}}{T_i^A} \alpha_r) - b^A B_i^A \right]^2}{T_i^A + T_i^{A2} \sigma_b^2 + B_i^A}$$

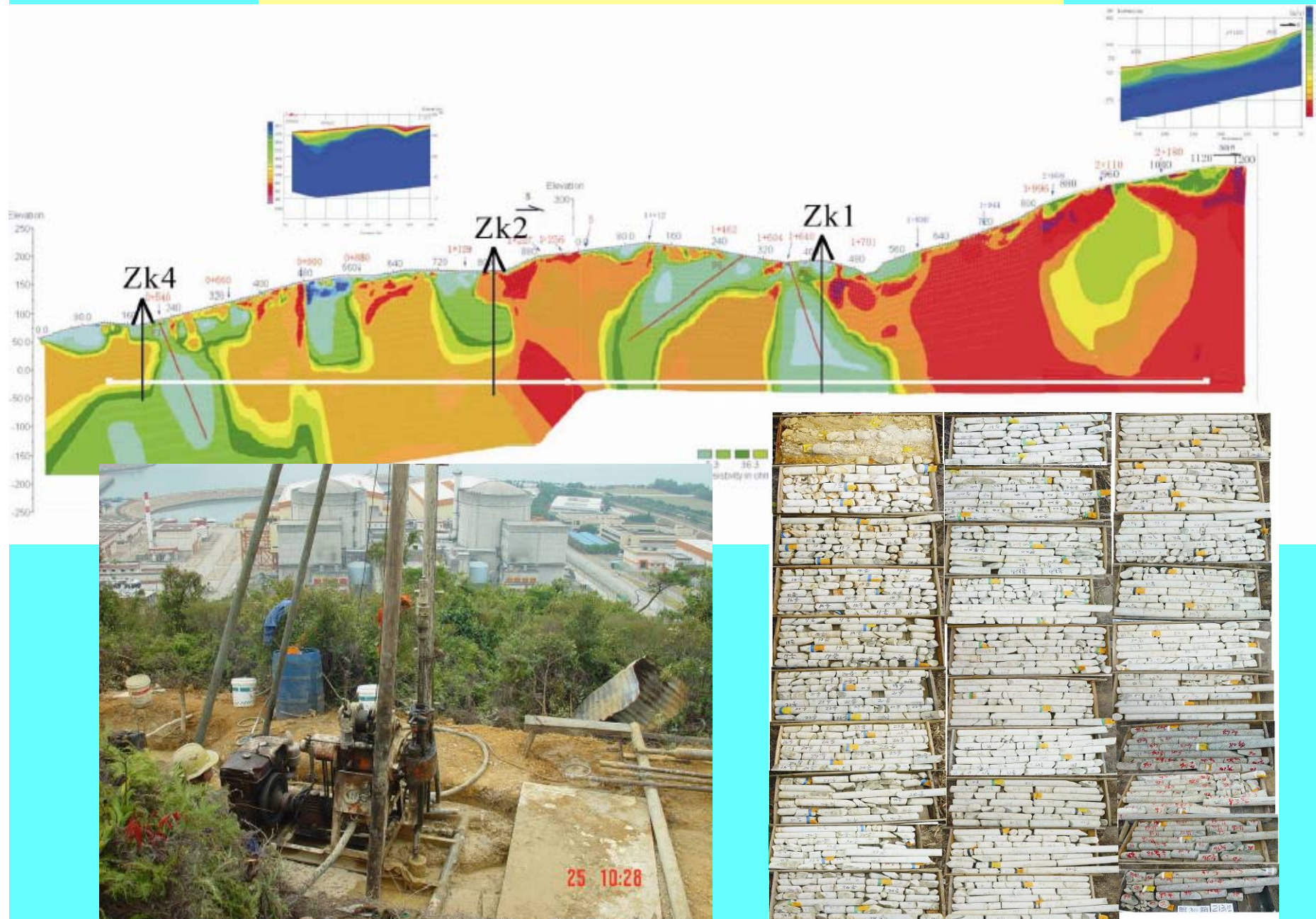
$$+ \frac{\alpha_D^2}{\sigma_D^2} + \frac{\alpha_c^2}{\sigma_c^2} + \sum_r \frac{\alpha_r^2}{\sigma_r^2} + \sum_{i=1}^{Nbin} \frac{c_i^2}{\sigma_{\text{shape}}^2} + \sum_{A=1,3} \left(\frac{\alpha_d^{A2}}{\sigma_d^2} + \frac{b^{A2}}{\sigma_B^2} \right)$$

Sensitivity to $\sin^2 2\theta_{13}$

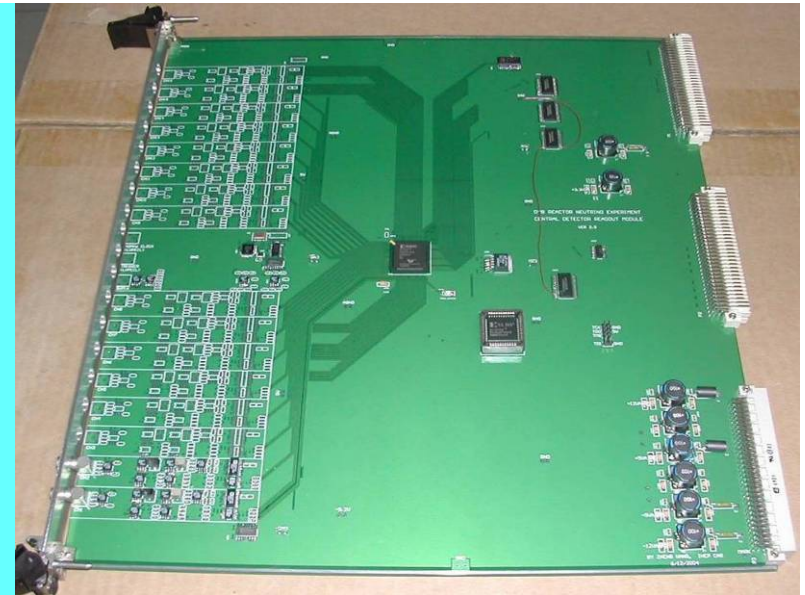


Other physics capabilities:
Supernova watch, Sterile neutrinos, ...

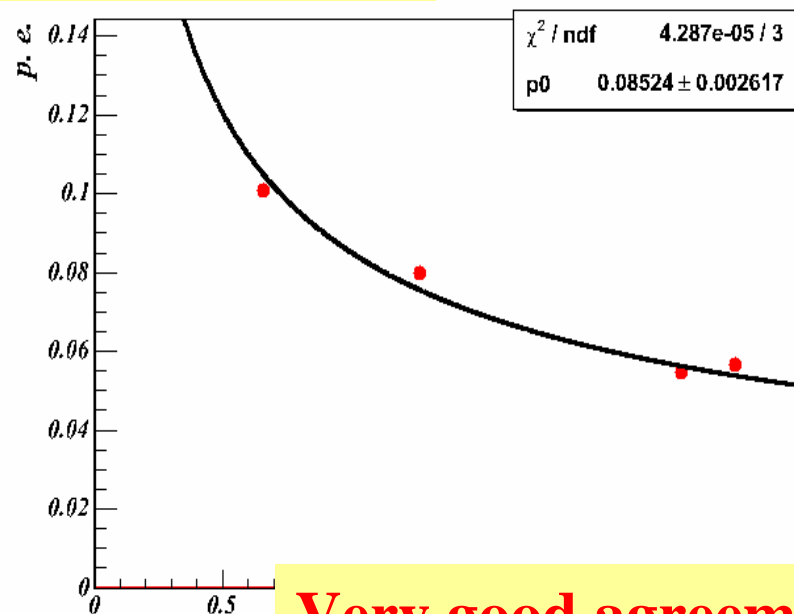
Site investigation completed



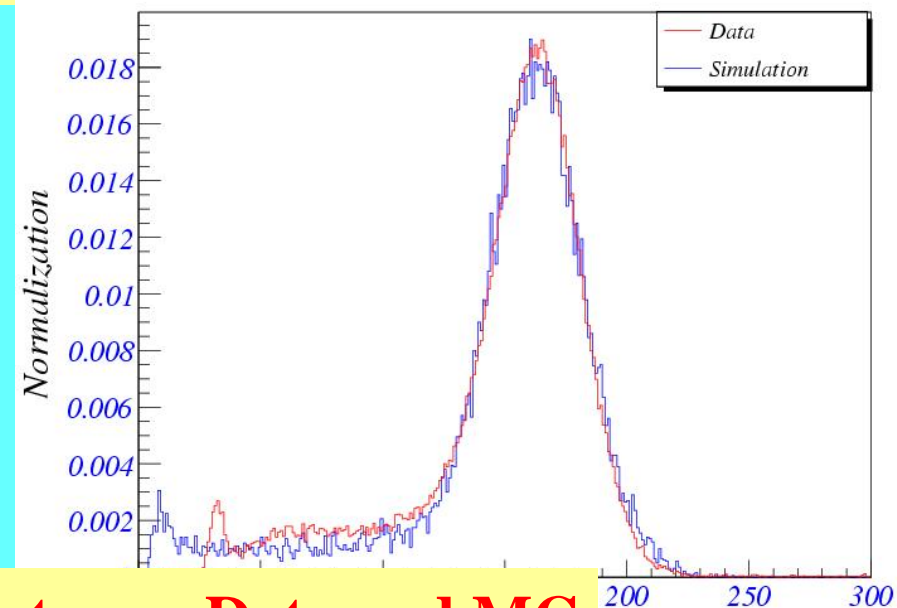
**Daya Bay Prototype:
45 PMT for 0.6 t LS**



Resolution: $8.5\%/\sqrt{E}$

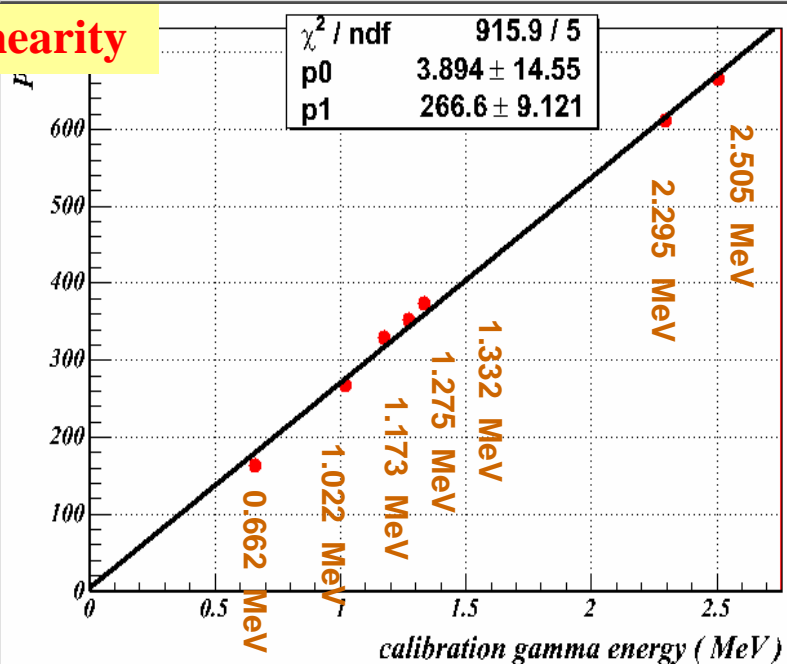


^{137}Cs spectrum

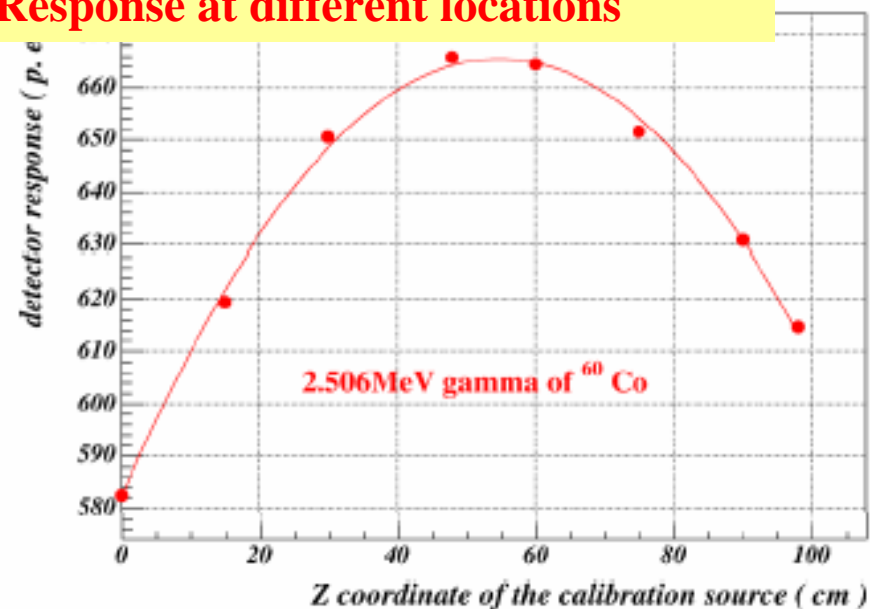


Very good agreement between Data and MC

linearity



Response at different locations



Daya Bay collaboration

Political Map of the World, June 1999



Status of the project

- CAS officially approved the project
- Chinese Atomic Energy Agency and the Daya Bay nuclear power plant are very supportive to the project
- Funding agencies in China are supportive, R&D funding in China approved and available
- R&D funding from DOE approved
- Site survey including bore holes completed
- R&D started in collaborating institutions, the prototype is operational
- Proposals to governments under preparation
- Good collaboration among China, US and other countries

Schedule of the project

- **Schedule**
 - **2004-2007** **R&D, engineering design, proposals for funding**
 - **2007-2008** **Civil Construction**
 - **2007-2009** **Detector construction**
 - **2009-2010** **Installation and testing
(one Near hall running)**
 - **2010** **Begin operations with full detector**

Daya Bay Collaboration Meeting

IHEP, Beijing, Feb. 13-15, 2006

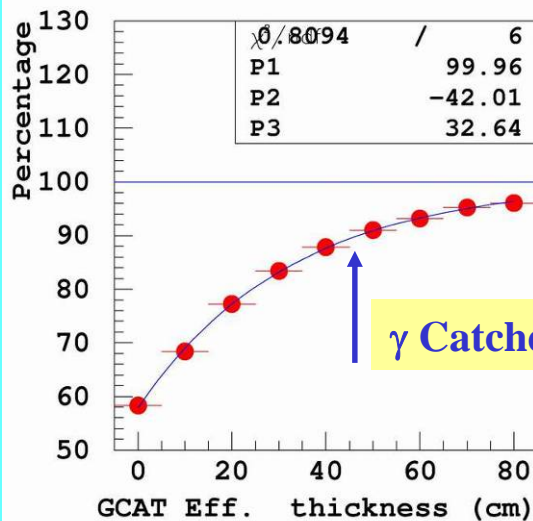
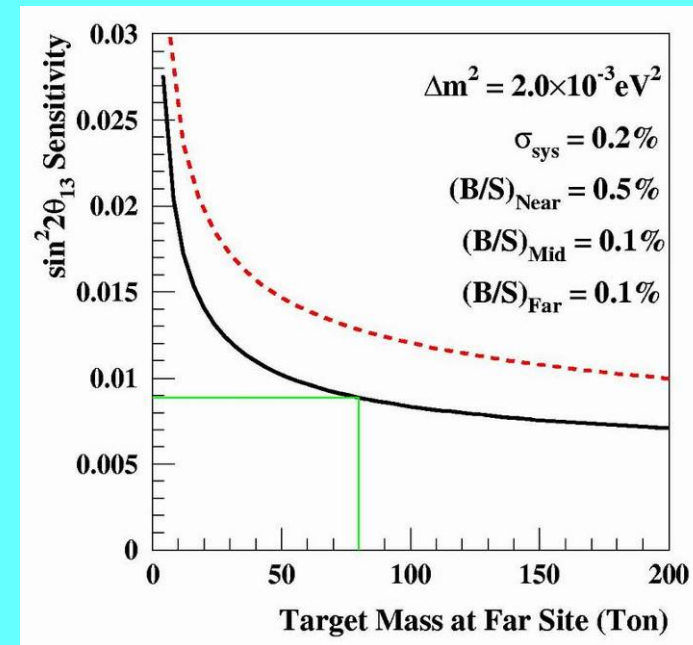


Detector dimension

Target mass: 20 t

Dimension of target:

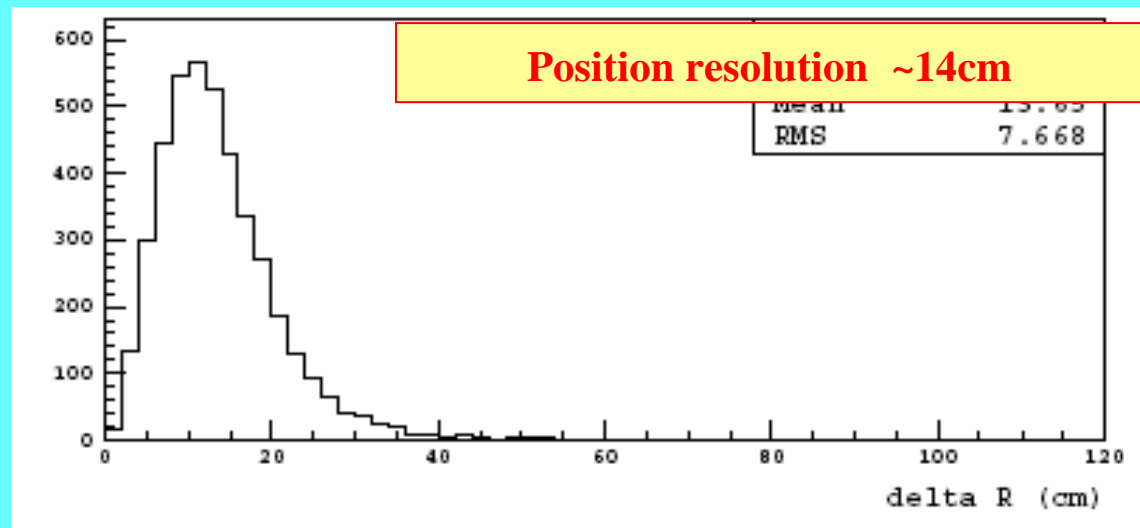
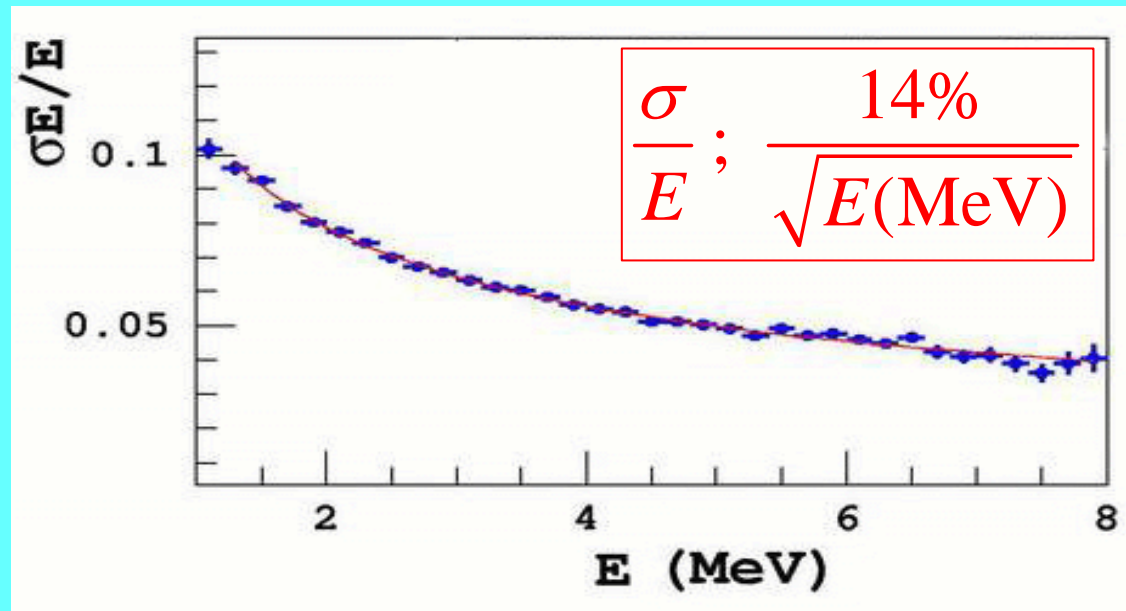
3.2 m × 3.2m



Oil buffer thickness

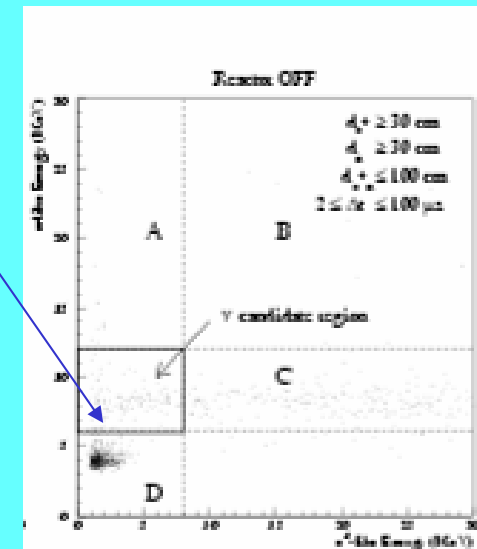
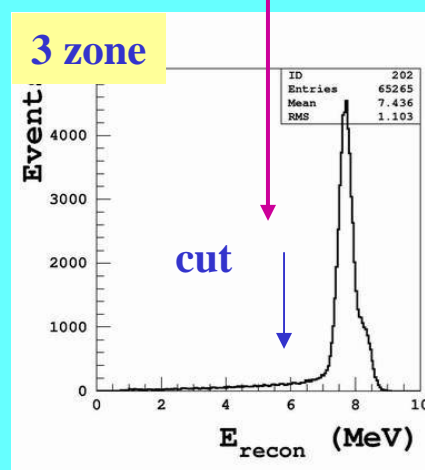
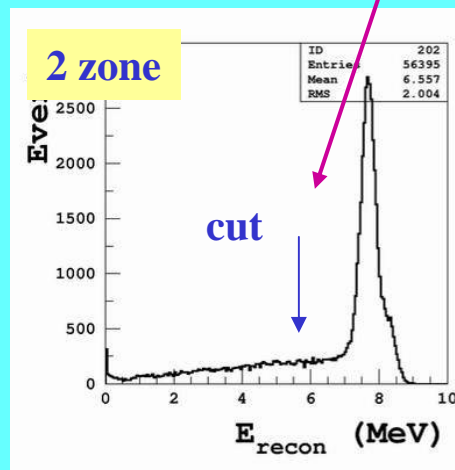
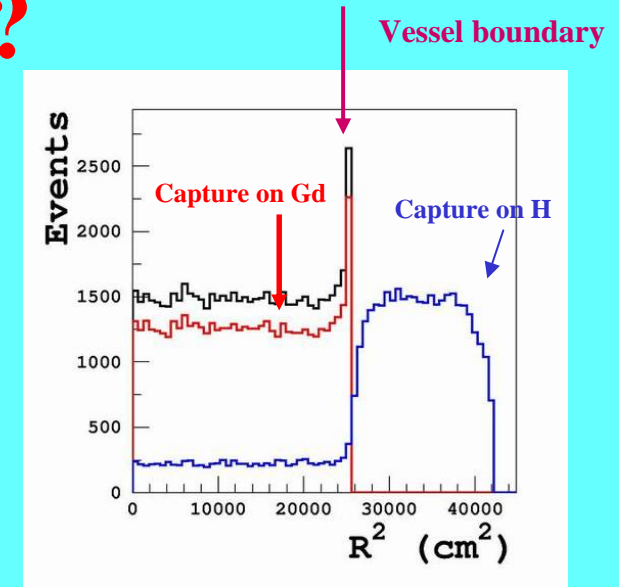
Isotopes	Purity (ppb)	20cm (Hz)	25cm (Hz)	30cm (Hz)	40cm (Hz)
$^{238}\text{U}(>1\text{MeV})$	50	2.7	2.0	1.4	0.8
$^{232}\text{Th}(>1\text{MeV})$	50	1.2	0.9	0.7	0.4
$^{40}\text{K}(>1\text{MeV})$	10	1.8	1.3	0.9	0.5
Total		5.7	4.2	3.0	1.7

Resolution



Why three zones ?

- Three zones:
 - Complicated acrylic tank construction
 - γ backgrounds on walls
 - Less fiducial volume
- Two zones:
 - Neutrino energy spectrum distorted
 - Neutron efficiency error due to energy scale and resolution:
 - two zones: 0.4%, three zones 0.2%
 - Using 4 MeV cut can reduce the error by a factor of two, but backgrounds from $\beta+\gamma$ do not allow us to do so



Reactor-related Uncertainties of Daya Bay

- The error due to power fluctuations of the reactors is given by:

$$\sigma_{sys} = \sigma_p \sqrt{\sum_r (f_F^r - f_N^r)^2}$$

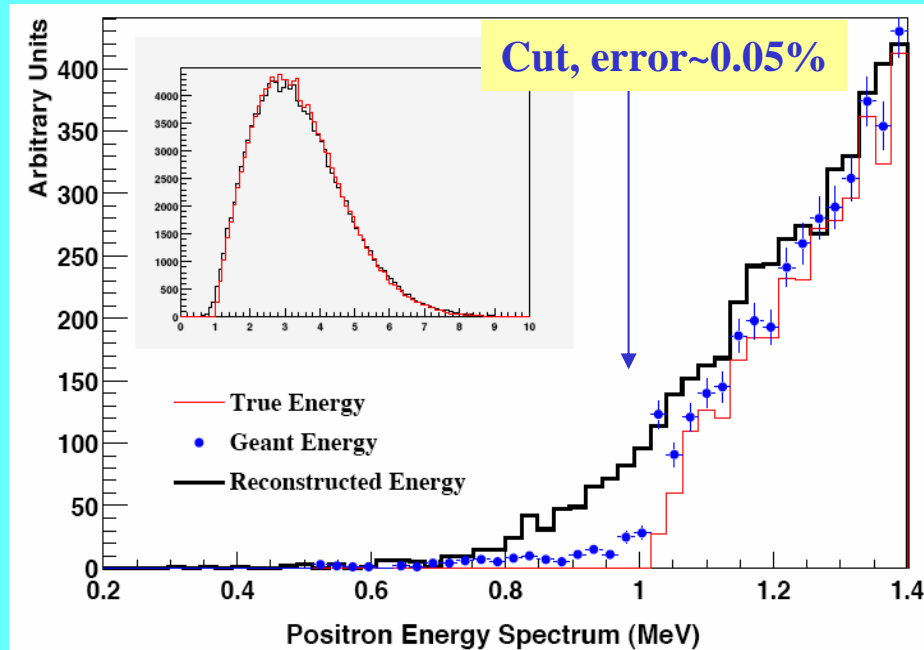
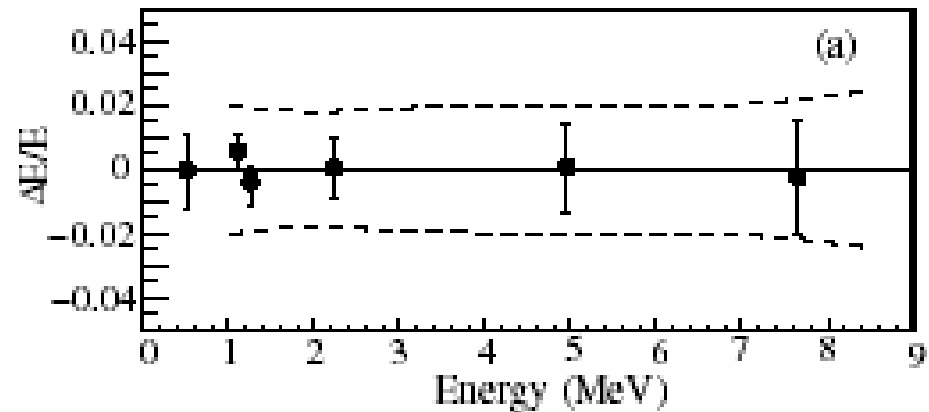
Based on experience of past experiments, due to uncertainty in measuring the amount of thermal power produced, the uncorrelated error per reactor core $\sigma_p \approx 2\%$.

f_F^r and f_N^r are fractions of the events at the far and near site from reactor r respectively.

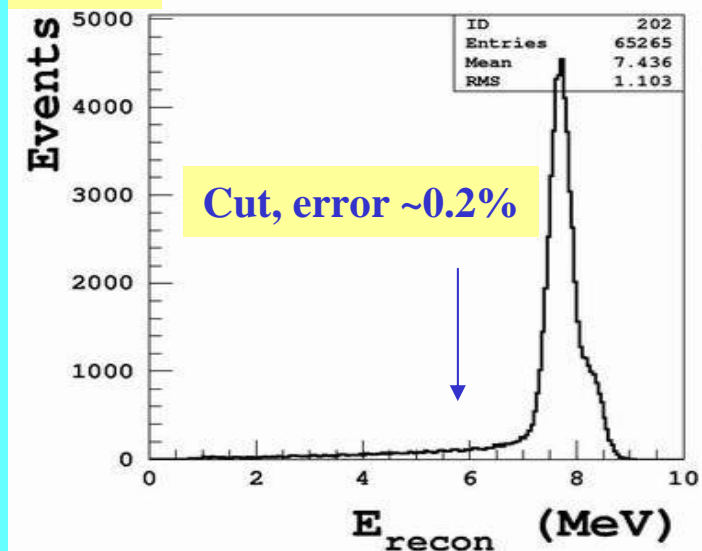
# Reactor Cores	Syst. error due to Power Fluctuations	Syst. error due to Core Positions	Total syst. error
4	0.035%	0.08%	0.087%
6	0.097%	0.08%	0.126%

Energy Cuts

Dominated by energy scale
KamLAND ~ 1%



3 zone

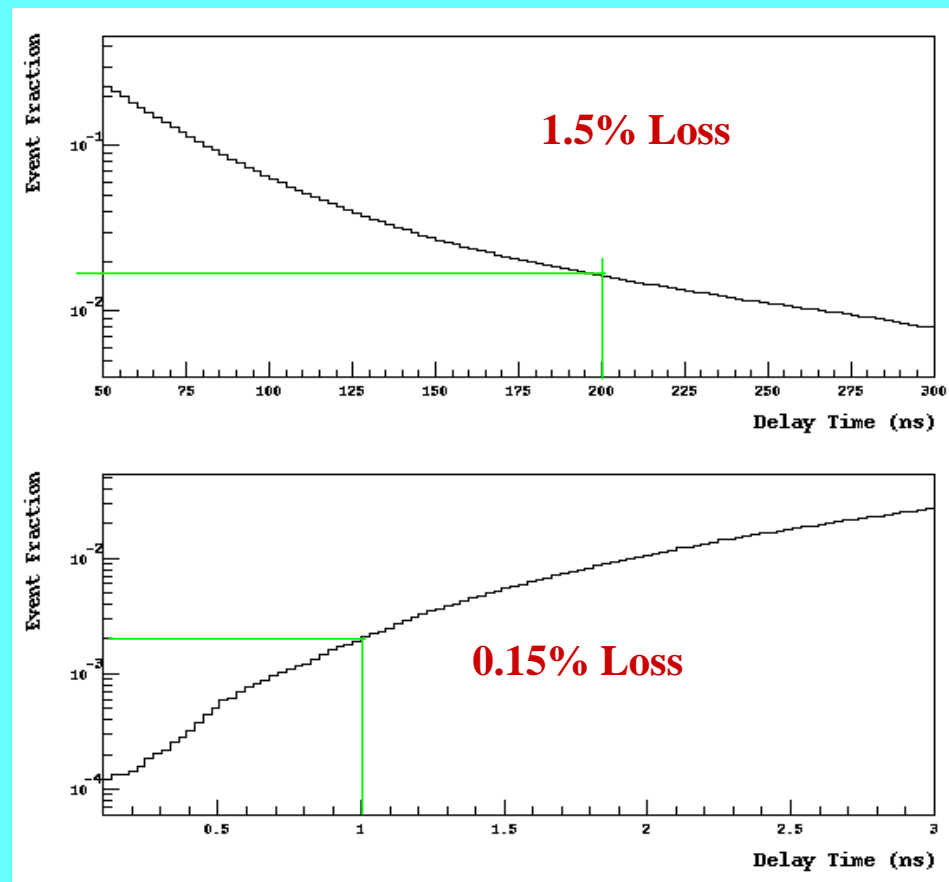


Time Cuts

Neutron time window uncertainty:

- $\Delta t = 10 \text{ ns} \rightarrow 0.03\%$ uncertainty
- Use common clock

→ Baseline = 0.1%
→ Goal = 0.03%



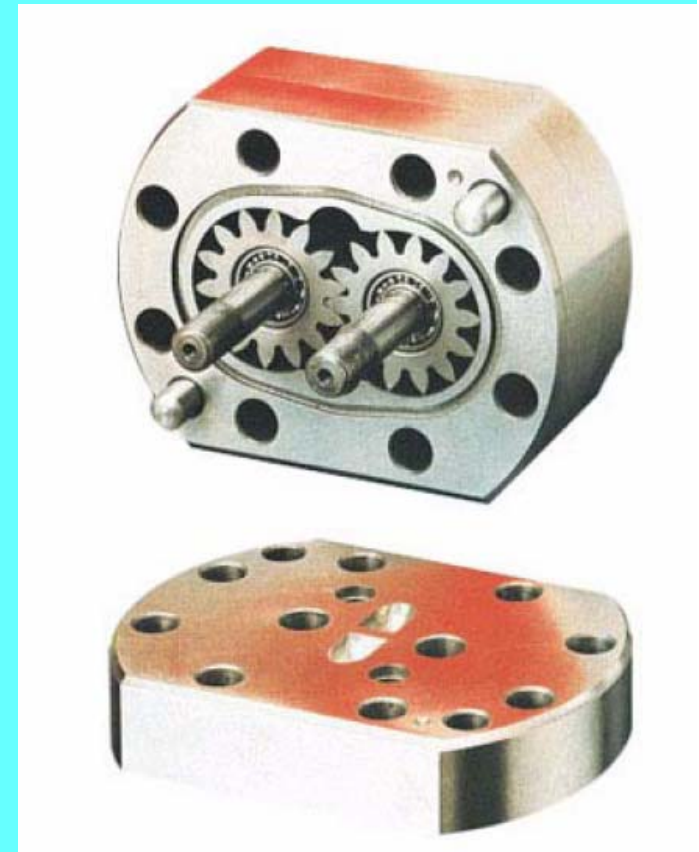
Livetime

- Measure relative livetimes using accurate common clock
- Use LED to simulate neutrino events
- Should be negligible error

Target Volume

Can be cancelled by swapping

- KamLAND: ~1%
 - CHOOZ: 0.02%?
 - Flowmeters - 0.02% repeatability
- Baseline = 0.2%
- Goal = 0.02%



H/C and H/Gd ratio

Can be cancelled by swapping

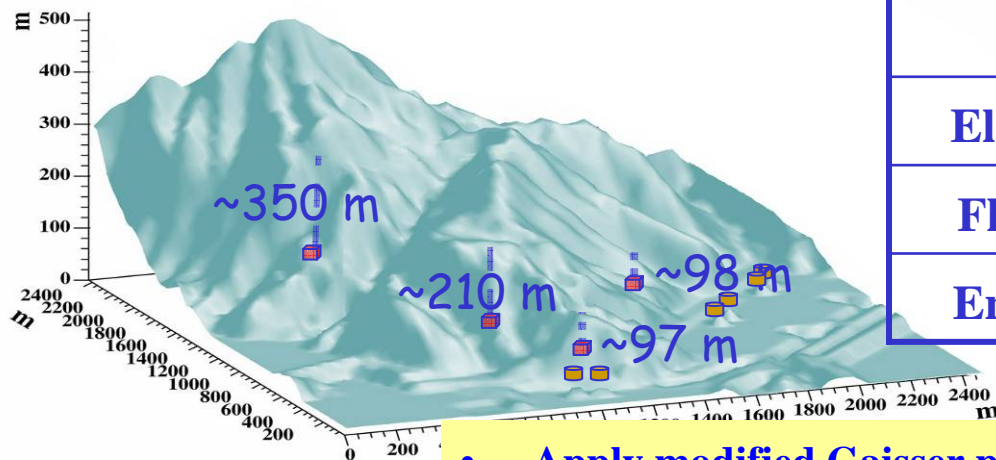
- H/C ratio

- CHOOZ claims 0.8% absolute based on multiple lab analyses (combustion)
- Use well defined liquid such as LAB and dodecane
- R&D: measure via NMR or neutron capture
- Expected error: 0.1%-0.2%

- H/Gd ratio

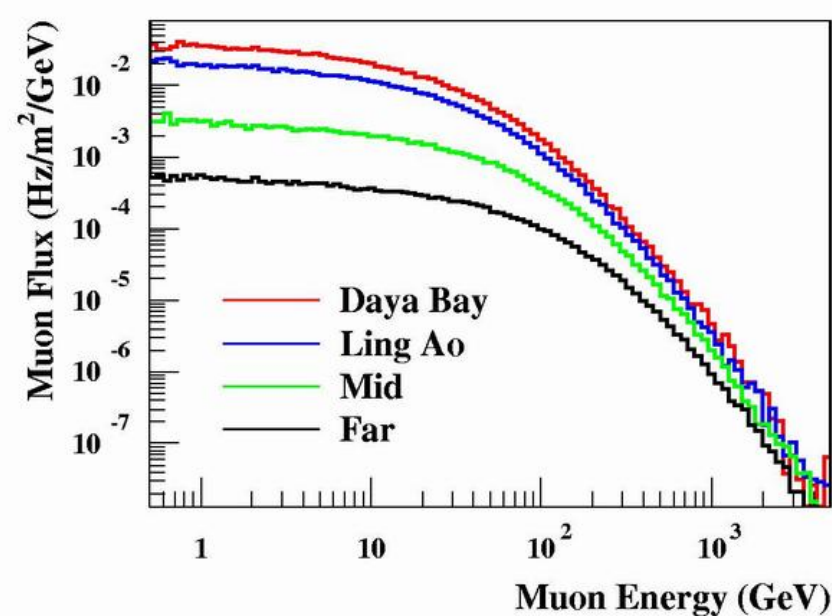
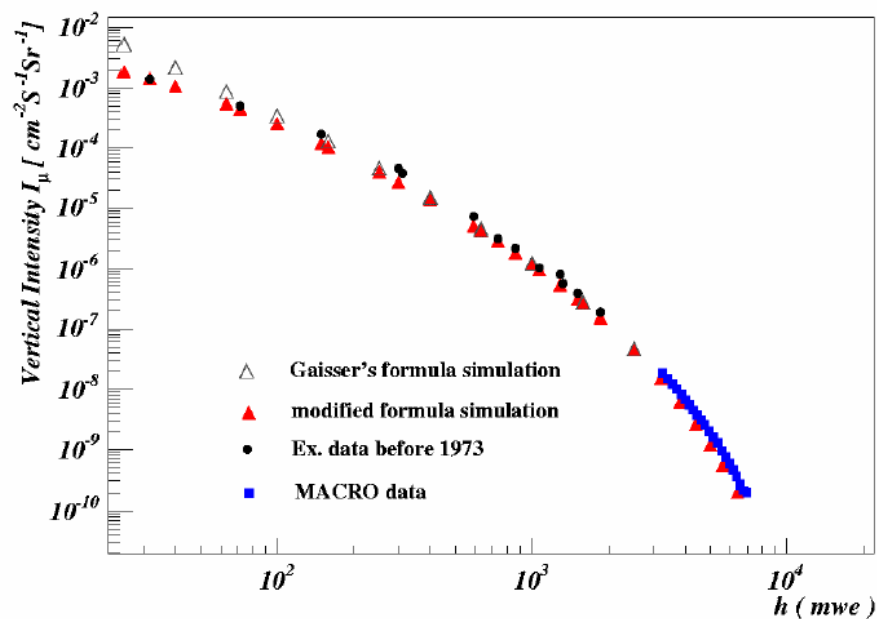
- Can be measured by neutron activated x-rays and neutron capture time
- For $\Delta t = 0.5 \mu\text{s}$, \rightarrow error $\sim 0.02\%$

Cosmic-muons at the laboratory

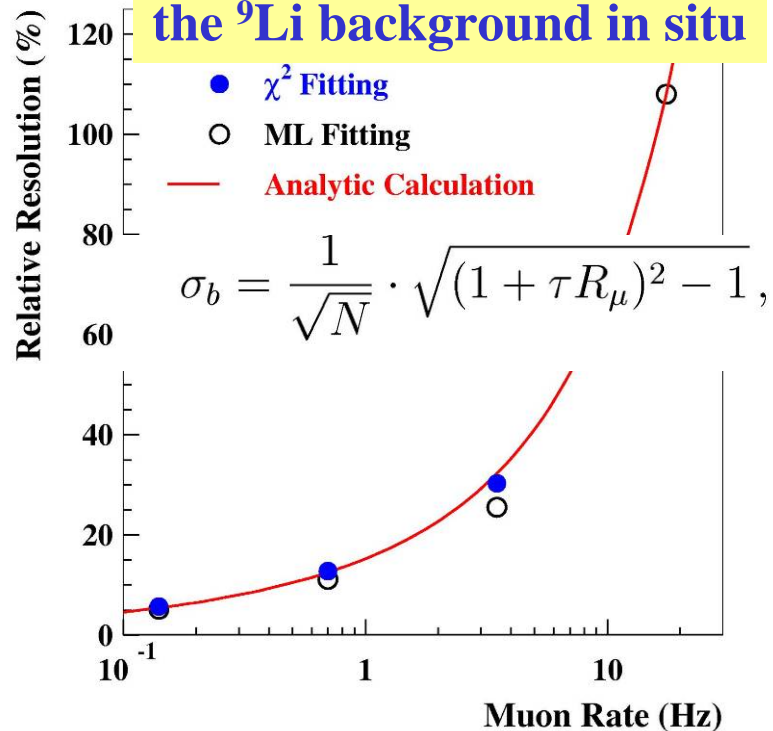


	DYB	LA	Mid	Far
Elevation (m)	97	98	208	347
Flux (Hz/m ²)	1.2	0.73	0.17	0.045
Energy (GeV)	55	60	97	136

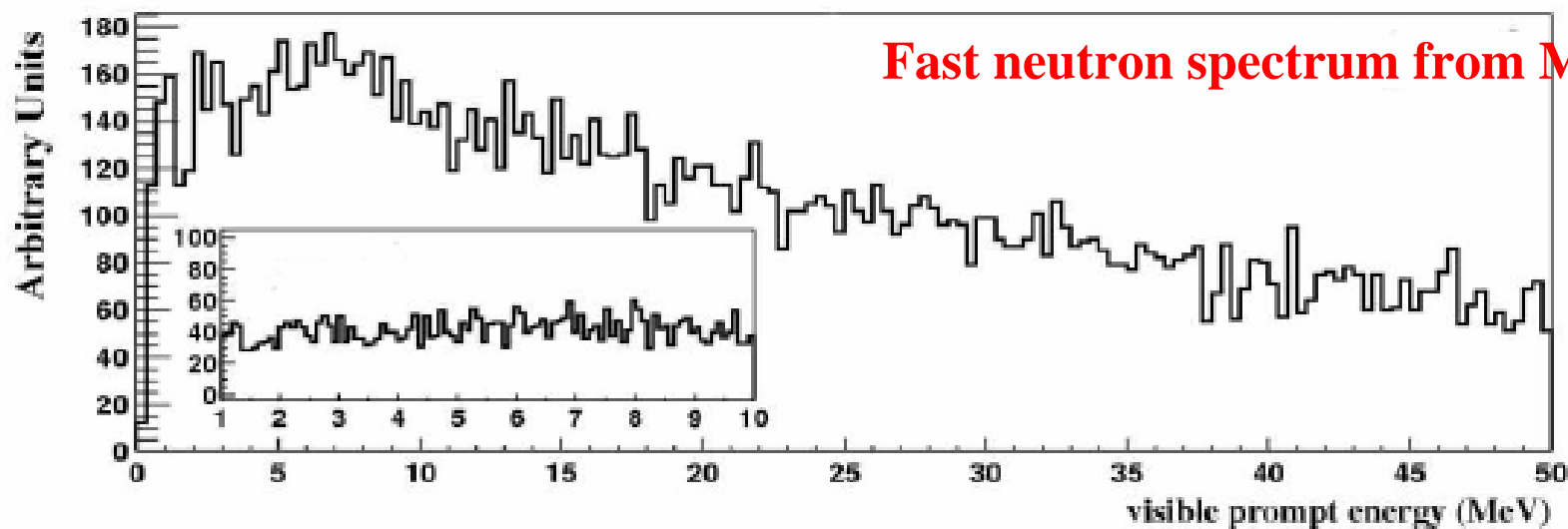
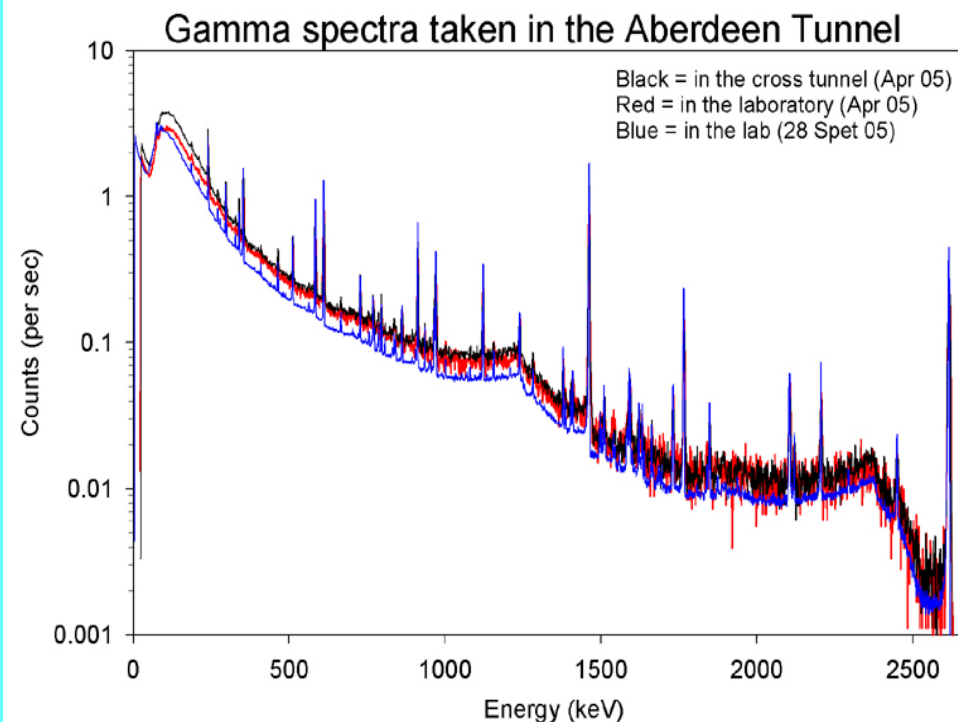
- Apply modified Gaisser parametrization for cosmic-ray flux at surface
- Use MUSIC and mountain profile to estimate muon flux & energy



Precision to determine the ^9Li background in situ



Spectrum of accidental background



Tunnel construction

- The tunnel length is about 3000m
- Local railway construction company has a lot of experience (similar cross section)
- Cost estimate by professionals, ~ 3K \$/m
- Construction time is ~ 15-24 months
- A similar tunnel on site as a reference

